

Moon Calculator

Version 6.0

Program and documentation by

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"The sun must not catch up the moon, nor does the night outstrip the day. Each one is travelling in an orbit with its own motion"
(Al Qur'an 36:40)

"the sun and the moon (are subjected) to calculations"
(Al Qur'an 55:5)

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(*as appropriate)

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<http://www.starlight.demon.co.uk/mooncalc>
<http://www.ummah.org.uk/ildl/mooncalc.html>

MoonCalc may be copied and distributed freely as long as all files are copied and no charge is made (other than a nominal charge for media). The program must be distributed as its ORIGINAL and UNMODIFIED zip file moonc60.zip).

No alterations should be made to the program, documentation or data files apart from the atlas database, TOWNS.DAT.

Although MoonCalc may be distributed freely, it is not 'Public Domain' nor is it 'Freeware'. All rights remain with the author, Dr. Monzur Ahmed.

1. Introduction

MoonCalc provides information relating to the position, age, phase, orientation, appearance and visibility of the moon for any given date, time and location on earth. It also provides the Julian Day Number, Magnetic Declination, time and direction of moonrise and moonset, interval between sunset and moonset, interval between sunrise and moonrise, date/time of astronomical new moon (conjunction), full moon, apogee and perigee and predicts the likelihood of visualising the young moon from a particular location. Data pertaining to solar and lunar eclipses in any year are also shown. MoonCalc provides Hijri calendar data including location dependent Hijri date conversion using predicted crescent visibility. Automatic local and regional (tri-zonal) Hijri calendar tabulation is possible.

The program can scan the globe at the start of any lunar month to find the location, date/time and direction of earliest crescent sighting using a variety of ancient and modern moon sighting criteria. The program is able to draw world maps (flat and spherical projections) showing areas of the globe where the young moon is likely to be seen.

Graphical displays showing the position of the moon on a star chart and the position of the moon in a simulated local sky (horizon view or traditional circular sky-chart view) can be produced and printed out. A close-up of the near side of the moon (showing orientation of the moon's limbs and position of the lunar craters), correct for a given observation site, is also provided. This close up takes into account the effect of libration and 'limb shortening' (optional). A graph of lunar libration for an entire month can be plotted.

There is a choice of either topocentric/geocentric co-ordinates and apparent/geometric sunset. Correction for atmospheric refraction is optional.

The program has a built in atlas database which stores latitude and longitude data of upto 1000 cities (ships with over 100 cities already entered). There are many user-configurable features.

2. Getting Started

2.1 Minimum system requirements

The minimum system requirements needed to run MoonCalc are:

- 386 based PC or compatible running DOS (486DX or better recommended)
- One floppy drive (hard drive strongly recommended)
- At least 500K free on disk for temporary storage (floppy should not be write protected)
- Colour VGA display or better (partial support for CGA, EGA and Hercules displays)

2.2 Files included

The following 14 files should be included on the distribution disk or be present after unzipping the moonc60.zip archive:

MOONC60.EXE	The main program
DEFAULTS.MC	Stores initial program default values
TOWNS.DAT	Database of town data, can be modified*
STARDATA.DAT	Data of 9025 stars from Yale Brightstar Database
BRIGHTST.DAT	Data of 1st magnitude stars
MOONFACE.DAT	Data to generate lunar craters
CONSTELN.DAT	Data to generate constellation lines
WORLDMAP.DAT	Data to generate world map
MAGMODEL.DAT	Model data for calculating Magnetic Declination
SVGA256.BGI	Required to display 600x800 and 1024x768 graphics modes (BGI driver copyright Borland Intl.)
MOONC60.BMP	Nice icon for use as windows shortcut
README.TXT	Documentation for MoonCalc in ASCII format
README.PDF	This file!
WHATSNEW.TXT	Lists new features + history of release dates.

The following file is generated by the program:

SCAN.DAT	Temporary file produced by program during scanning.
----------	---

(* in versions prior to 5.0, this file was called DATA.PTC and had a different structure)

2.3 Making Backups

As with all new programs, it is advisable to make backup copies of all the files. You should then write protect the original disk and keep it in a safe place. Use only the backed up disk.

2.4 Running MoonCalc on a floppy drive system

DOS

Place your disk in, say, drive A. Now make sure you have the A:> prompt showing:

A:>

Type MOONC60 <CR> and the program will start with standard graphics mode.

MOONC60 S <CR> will start program with graphics in 600x800 mode.

MOONC60 H <CR> will start program with graphics in 1024x768 mode.

The floppy should NOT be write protected as MoonCalc will need to access the disk for temporary storage.

Windows

Place your disk in, say, drive A. Use Windows explorer or File Manager to log onto drive A. Then double-click on the MOONC60.EXE file.

2.5 Installing and running MoonCalc on a hard drive system

DOS

Let us assume that your hard drive is called drive C. You should initially make a directory called e.g. MOON:

```
md c:\moon <CR>
```

Put the floppy disc containing the program into drive A. Copy all the files from the floppy disc into the MOON directory:

```
copy a:*. * c:\moon <CR>
```

Ensure that you are logged onto the MOON directory:

```
cd c:\moon <CR>
```

Now type MOONC60 <CR> and the program will start with standard graphics mode.

MOONC60 S <CR> will start program with graphics in 600x800 mode.

MOONC60 H <CR> will start program with graphics in 1024x768 mode.

Windows

MoonCalc may be run in a DOS box under Windows 3.1 or Windows 95/98. This is the preferred option. You can have a desktop shortcut if you wish.

When using MoonCalc under Windows make sure that the 'Working' or 'Start in' property of the desktop shortcut points to the directory that contains the MoonCalc files. This tells MoonCalc where to find the associated files.

The 'Cmd line' property of the shortcut should be...

MOONC60 to start program with standard graphics mode (640x480).

MOONC60 S to start program with graphics in 600x800 mode.

MOONC60 H to start program with graphics in 1024x768 mode.

The latter two options may not work on all systems and I do not recommend the last one unless you have a very good monitor!

3.1.5 Earliest new moon sighting for a given location

To find the date and time of earliest sighting of the new moon for a given location, use the following steps:

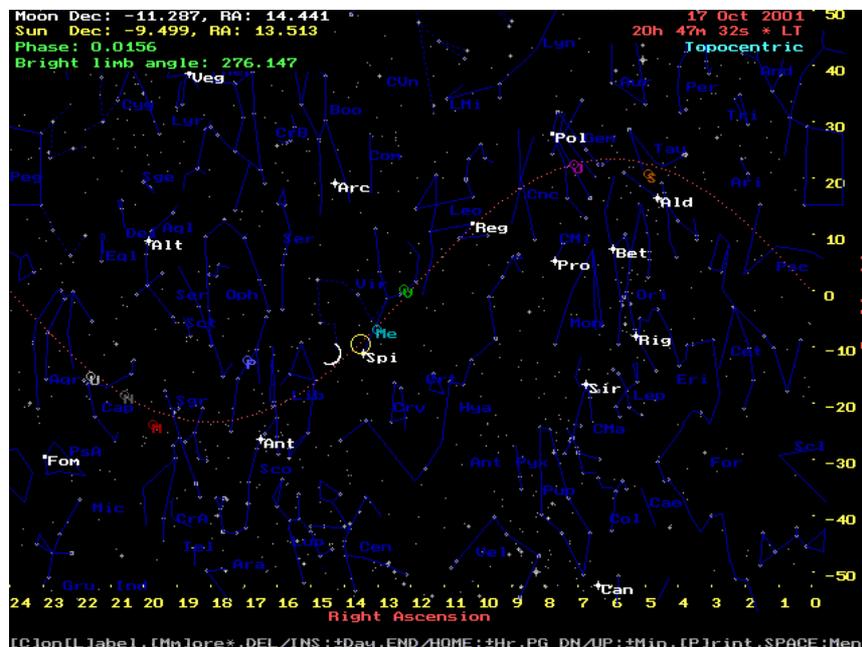
- Choose option 1 from Main Menu and enter the name and details of the location in question.
- Enter a date near the time of interest and obtain the date of the astronomical new moon (also known as date of conjunction).
- Go back to the Main Menu, choose option 1 again and enter this date.
- Go to screen 2 of 4 which shows the data at local sunset.
- The program will probably say that moon is 'Not visible' or 'Moon not new' ('Moon not new' means that the moon is over 7 days old and implies that it should be visible).
- Use the DEL/INS key to 'hunt' around this date until the earliest date when moon 'Should be visible' is obtained.
- Now go to screen 3 of 4 to obtain the data for when the sun is 5 degrees below the horizon i.e. the optimum time of sighting, azimuth etc.
- To get the information for the following month, go back to Screen 2 of 4 and press the '+' key to jump forward one lunar month. Again 'hunt' with the DEL/INS keys to obtain date of earliest sighting. Now go to screen 3 of 4 as before.

It may sound complicated but after a while you will find the procedure straightforward.

Alternatively, just go straight to screen 4 of 4 which performs the above steps automatically and shows the date when the previous new crescent would have been seen for the first time from you location (using the visibility criterion that you have chosen).

3.2 Option 2. Moon position on Starchart (Dec versus RA)

When this option is chosen, you will be required to enter the date and time. This option produces a star chart (a graph of Declination Angle versus Right Ascension) and plots the positions of the moon, sun and planets on it. The ecliptic is shown as a red line.



Starchart: yellow circle=Sun; white disc/crescent= Moon, Me= Mercury, V=Venus, M=Mars, J=Jupiter, S=Saturn, U=Uranus, N=Neptune, P=Pluto; stars and constellations – see text.

The positional data for the stars were obtained from the Yale Brightstar Database. The first magnitude stars are labelled using the first 3 letters of their common names. NB Pol is Pollux not Polaris (Pole Star)!

COMMON NAME MAGNITUDE

Sirius	-1.46
Canopus	-0.72
Alpha Centuri	-0.27
Arcturus	-0.04
Vega	0.03
Capella	0.08
Rigel	0.12
Procyon	0.38
Achernar	0.46
Betelgeuse	0.50
Hadar	0.61
Altair	0.77
Acrux	0.83
Aldebaran	0.85
Antares	0.96
Spica	0.98
Pollux	1.14
Fomalhaut	1.16
Deneb	1.25
Becrux	1.25
Regulus	1.35

The planetary positions are also plotted and labelled:

Me	Mercury
V	Venus
M	Mars
J	Jupiter
S	Saturn
U	Uranus
N	Neptune
P	Pluto [position not accurate for dates before 1890 CE]

You can view and label the constellations. The constellation lines are based on information in Patrick Moore's Guinness Book of Astronomy (4th edition) (ref 34) which in turn refers to the Cambridge Sky Catalogues (1987). The 3 letter abbreviations for the constellations are as follows:

ANDROMEDA = And	ANTLIA = Ant	APUS = Aps
AQUARIUS = Aqr	AQUILA = Aql	ARA = Ara
ARIES = Ari	AURIGA = Aur	BOOTES = Boo
CAELUM = Cae	CAMELOPARDUS = Cam	CANCER = Cnc
CANES VENATICI CVn	CANIS MAJOR = CMa	CANIS MINOR = CMi
CAPRICORN = Cap	CARINA = Car	CASSIOPEIA = Cas
CENTAURUS = Cen	CEPHEUS = Cep	CETUS = Cet
CHAMAELEON = Cha	CIRCINUS = Cir	COLUMBA = Col
COMA BERENICES = Com	CORONA AUSTRALIS = CrA	CORONA BOREALIS = CrB
CORVUS = Crv	CRATER = Crt	CRUX AUSTRALIS = Cru
CYGNUS = Cyg	DELPHINUS = Del	DORADO = Dor
DRACO = Dra	EQUULEUS = EqL	ERIDANUS = Eri
FORNAX = For	GEMINI = Gem	GRUS = Gru
HERCULES = Her	HOROLOGIUM = Hor	HYDRA = Hya
HYDRUS = Hyi	INDUS = Ind	LACERTA = Lac
LEO = Leo	LEO MINOR = Lmi	LEPUS = Lep
LIBRA = Lib	LUPUS = Lup	LYNX = Lyn
LYRA = Lyr	MENSA = Men	MICROSCOPIUM = Mic
MONOCEROS = Mon	MUSCA AUSTRALIS = Mus	NORMA = Nor
OCTANS = Oct	OPHIUCHUS = Oph	ORION = Ori
PAVO = Pav	PEGASUS = Peg	PERSEUS = Per
PHOENIX = Phe	PICTOR = Pic	PISCES = Psc

PISCIS AUSTRALIS = PsA	PUPPIS = Pup	PYXIS = Pyx
RETICULUM = Ret	SAGITTA = Sge	SAGITTARIUS = Sgr
SCORPIUS = Sco	SCULPTOR = Scl	SCUTUM = Sct
SERPENS = Ser	SEXTANS = Sxt	TAURAS = Tau
TELESCOPIUM = Tel	TRIANGULUM = Tri	TRIANGULUM AUSTRALE = TrA
TUCANA = Tuc	URSA MAJOR = Uma	URSA MINOR = UMi
VELA = Vel	VIRGO = Vir	VOLANS = Vol
VULPECULA = Vul		

The bright limb angle, phase, Right Ascension and Declination of the moon and Right Ascension and Declination of the sun are also depicted.

Once the stargram is drawn, the following keys apply:

C:	Draw/remove major constellation lines
L:	Label/unlabel major constellations with standard 3 letter abbreviation
P:	Print screen to Epson/HP compatible printer
M/m:	Show more/less stars (i.e. change magnitude)
SPACE:	Return to main menu

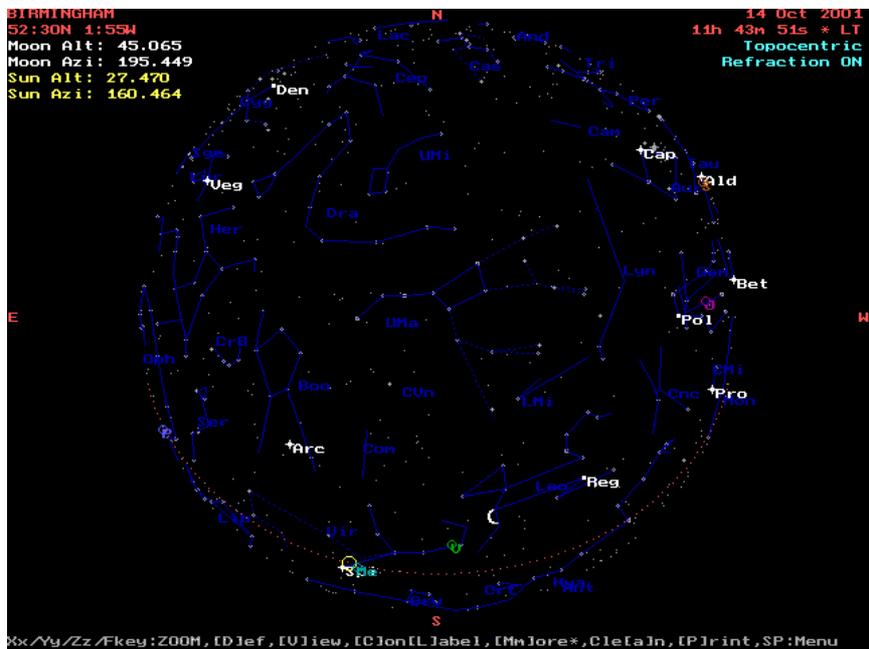
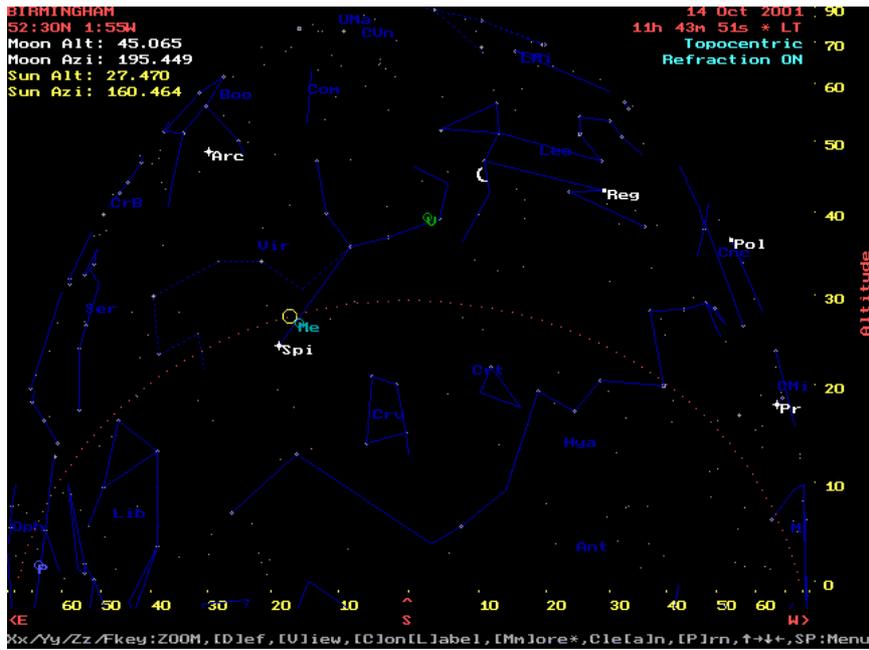
Also, as before...

+/-	increase/ decrease MONTH
DEL/INS	increase/ decrease DAY
END/HOME	increase/ decrease HOUR
PAGE DN/UP	increase/ decrease MINUTE

3.3 Option 3. Simulation of Local Sky (Alt versus Azi)

Enter the location and date/time data as usual. The maximum magnitude of stars to be displayed is also required. The star database in the program contains data for over 9000 stars down to magnitude seven (the lower the magnitude, the brighter the star). If you enter a high number for the maximum magnitude, the display will show more (and dimmer) stars but will take a long time to generate on a slower computer. On a slow machine it is better to view only the brighter stars (by specifying maximum magnitude as eg 2 or 3).

The program will now generate a simulation of the sky showing the position of the stars and planets. You can toggle between a 'horizon' view and a traditional 'circular' sky chart. The latter gives a view of the sky as it would appear if you were lying on your back with your head facing north and feet facing south:



Sky simulation: horizon view (above), circular view (below); Yellow circle=Sun; white disc/crescent= Moon, Me= Mercury, V=Venus, M=Mars, J=Jupiter, S=Saturn, U=Uranus, N=Neptune, P=Pluto; stars & constellations – see text

The positions of the moon and sun are drawn on this background of stars and planets. The moon is drawn in white showing the correct phase and orientation taking into account its bright limb angle and parallactic angle. The sun is represented by a yellow circle. A printout of the graphical screen can be made if an Epson dot matrix or HP Laserjet/Inkjet printer is connected.

Once the sky is drawn, the following keys apply:

- X/x, Y/y: Zoom in/out in the X and Y axes respectively
- Z/z: Zoom in/out maintaining current aspect ratio
- Function keys 1-10: 10 Pre-set zooms (zoom factor 1.1 to 3)
- D: Set initial default zoom of 1
- Cursors: Change direction of view (only in horizon view)
- N,E,S,W Change direction of view (only in horizon view)
- P: Print screen to Epson/HP compatible printer
- C: Draw/remove major constellation lines

L: Label/unlabel major constellations with standard 3 letter abbreviation
A: Toggles between clear screen & labelled screen
M/m: Show more/less stars (i.e. change magnitude)
V: Toggle between horizon view and circular sky view
SPACE: Return to main menu

Other keys which apply but which are not shown at the bottom of the screen are the usual:

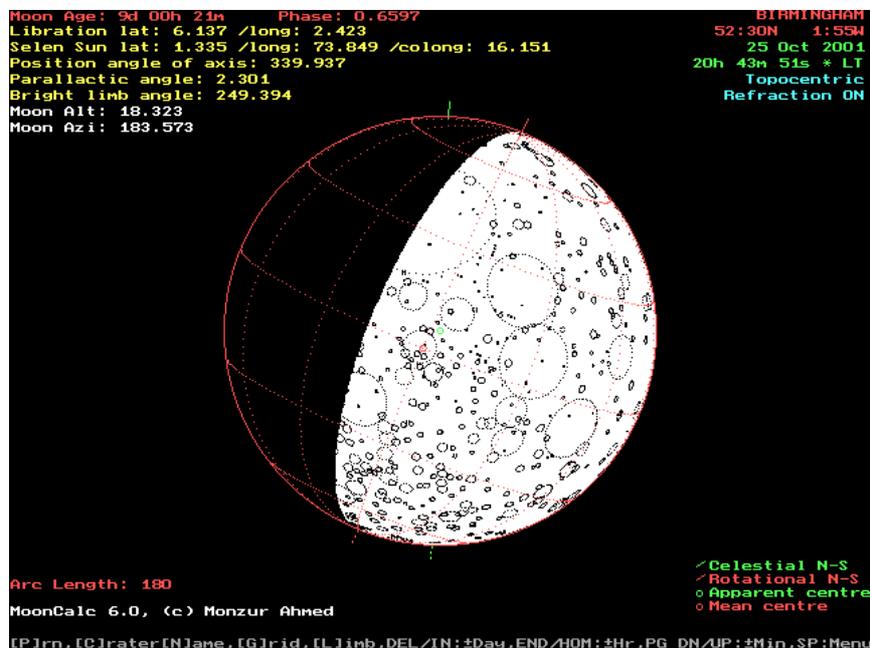
+/- increase/ decrease MONTH
DEL/INS increase/ decrease DAY
END/HOME increase/ decrease HOUR
PAGE DN/UP increase/ decrease MINUTE

We can 'see' the moon and sun setting or rising by increasing /decreasing the hour with END/HOME. We can zoom in and out of the central part of the sky using X/x, Y/y and Z/z. The function keys will produce images at preset zoom factors (Function key1=zoom factor 1.1, Function key 10= zoom factor 3).

3.4 Option 4. Close-up of Moon

As before, the place and date/time are first entered. A graphical representation of a close up of the moon is shown. The phase and orientation of the moon's limbs are depicted accurately.

The top left hand part of the screen also shows numeric values of phase, age, libration (latitude and longitude), position angle of axis, bright limb angle, parallactic angle, selenographic sun latitude, longitude and co-longitude, moon altitude and moon azimuth. The libration shown is the total (optical + physical) libration and is calculated using methods described by D.H. Eckhardt.



Close up of moon: craters can be switched on/off with 'C' and labelled/unlabelled with 'N'; grid can be switched on/off with 'G'; difference between apparent and mean centres reflects libration

Some of the above values will change slightly depending on whether MoonCalc is set to 'geocentric' or 'topocentric' and also if refraction is on/off (see sections 3.10.11 and 3.10.12). Please note that values of certain physical parameters of the moon given in the printed Almanac are geocentric. In particular, geocentric and topocentric libration may differ by as much as 1 degree. Topocentric reduction in the values for libration and position angle of axis are made using

differential corrections - equations in Explanatory Supplement to the Astronomical Ephemeris. See section 3.7 for more information on libration.

Again one can increase/decrease the month, day, hour or minute using the key combinations below and see the effect on the moon's appearance:

+/-	increase/ decrease MONTH
DEL/INS	increase/ decrease DAY
END/HOME	increase/ decrease HOUR
PAGE DN/UP	increase/ decrease MINUTE

This feature is especially useful for seeing how the orientation of the moon changes hour by hour.

Other keys which apply are:

'C': toggles between 'craters on' and 'craters off'. Switch on the craters feature if you have a fast machine (486DX or higher). When this feature is switched on, the craters and seas of the near side of the moon are depicted graphically taking libration into account.

'N': labels the larger craters/seas (remember 'N' for name)

'G': produces a latitude/longitude grid and shows the mean and apparent centres of the disc as well as the rotational and celestial axes.

'L': invokes 'limb shortening' i.e. for very thin crescents the tips of the crescent are not visible and so the crescent length is less than 180 degrees, sometimes considerably less. Pressing 'L' will not only shorten the crescent but will also display the approximate visible crescent length in degrees. MoonCalc uses algorithms developed from the data of Danjon (1932, 1936) and Schaefer (1991,ref 19) to shorten the crescent length.

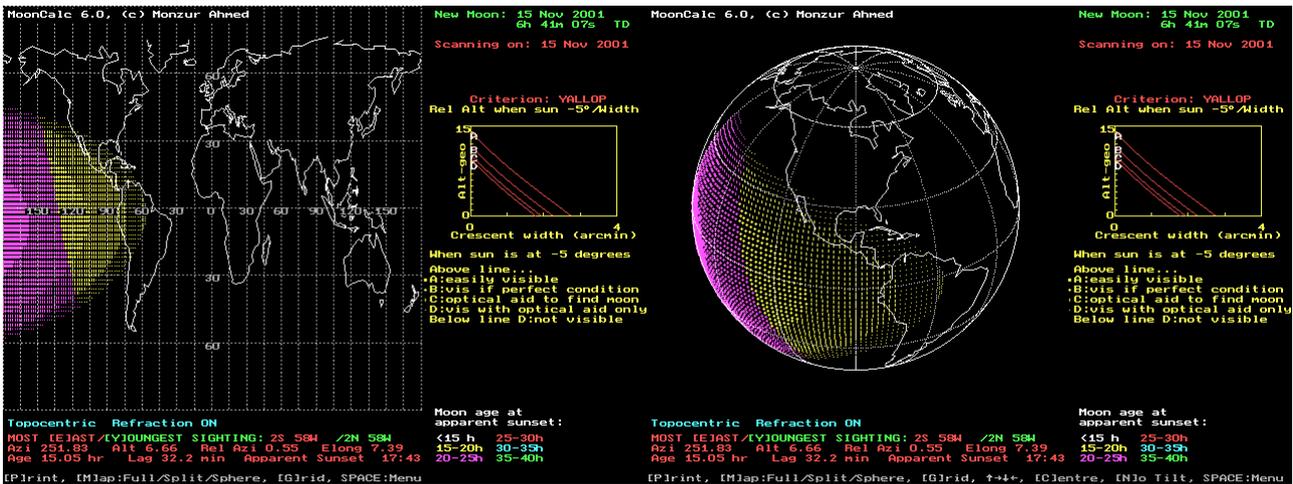
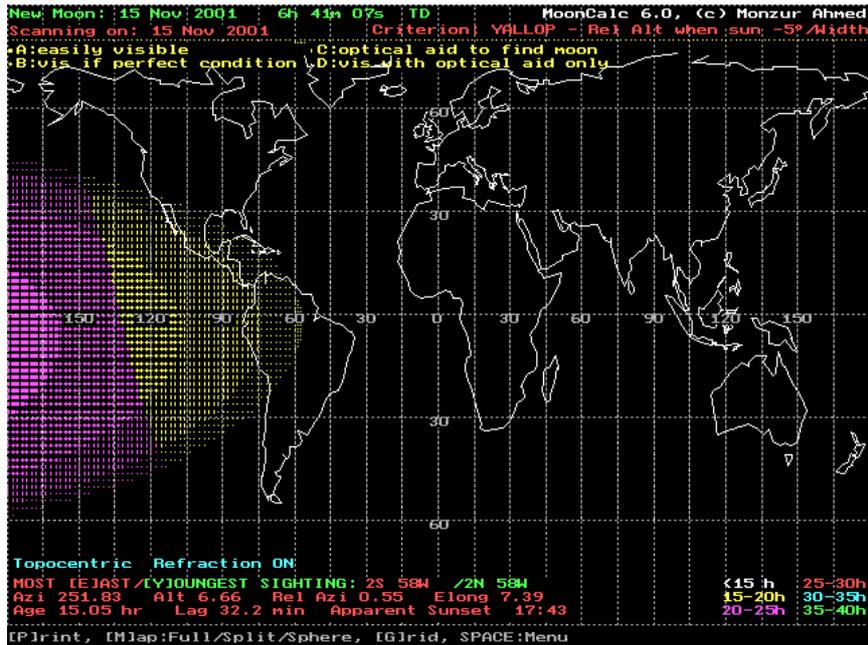
'P': a printout of the graphical screen can be made if an EPSON compatible or HP Laserjet/Inkjet printer is connected.

'SPACE': return to menu.

3.5 Option 5. First Crescent Sighting (Global Scan)

MoonCalc is able to predict the areas of the world where the young crescent moon is likely to be initially seen using one of several published/well known moon sighting criteria. The program will draw a world map and scan the world starting at longitude 180W and progress eastwards. The progress of the scan is indicated by a dotted yellow line near the top of the screen. The scan is performed in two passes (coarse scan first, then fine scan).

You can start scanning either on the day of conjunction or on the following day. At each longitude the program will search from a lower latitude (eg 60S) to a upper latitude (eg 60N). In other words, the world is divided into a fine grid and each intersection on the grid is examined to see if the new moon is visible at that location. If the minimum moon visibility criterion is satisfied by that location, then the location is marked with a coloured dot - the colour of the dot represents the age of the moon at local sunset (see lower right hand corner of the output screen for key to these colours). The map can be displayed in three ways:



Global scan: standard flat map (top), split screen with flat map (left), spherical map (right). Toggle between the three by pressing 'M'

At the end of the scan the program will display the location of the place where the moon will be first sighted (youngest sighting) as well as the most eastern location where the moon will be sighted (most easterly sighting). The location of the youngest sighting is usually slightly northwest or southwest of the most easterly point. The properties of the moon at the time of local sunset at the two points are shown at the bottom of the screen (use Y or E to toggle). After the scan is complete the following keys apply:

- P: printout of display
- M: change map layout (flat-full, flat-split, spherical-split)
- Cursors: used to spin spherical map
- N: remove tilt from spherical map (ie centre on 0 latitude)
- C: centre spherical map (ie centre on 0 longitude)
- G: show/hide latitude/longitude grid
- Y: show data for location of Youngest sighting.
- E: show data for most Easterly sighting.
- SPACE: return to menu

The global scan is a *very* processor intensive procedure and may take a long time to complete on slower computers. You can exit at any time during the scan by pressing ESC. If during a scan, you think that the scan has already located all possible areas where the new moon is likely to be seen you can save time by terminating the scan early by pressing any key (except ESC or SPACE). It is inadvisable to terminate prematurely if you are using the RGO 67 criterion as

- **Ilyas_A** [Alt, Elong]

Ilyas has written extensively on moon sighting and lunar calendars (eg ref 11,12,13 17 and 18). MoonCalc supports three of Ilyas' best known sighting criteria. The first criterion depends on the 'moon's relative altitude at sunset' and the 'sun-moon elongation at sunset' (ie angular separation between the sun and the moon). Again a curve based on observational data was drawn on a graph of moon's relative altitude at sunset versus sun-moon elongation at sunset. If the properties of a crescent lie above the curve then the crescent should be visible and vice versa.

- **Ilyas_B/modified Babylonian** [Lag, Latitude]

Ilyas' second criterion is a modification of the ancient Babylonian system of moonset lag times. However Ilyas compensates for latitude (eg at latitude 0 deg: lag 41 min; 30 deg:46 mins, 40 deg:49 mins, 50 deg: 55mins).

- **Ilyas_C** [Alt, Rel Azi]

Ilyas' third criterion, described in 1988, is a slight modification of Ilyas_A and depends on the moon's relative altitude at sunset and the difference in azimuth between the sun and moon at sunset.

- **RGO 67** [Alt, Elong]

The Royal Greenwich Observatory (which sadly closed recently) produced a series of information sheets which tabulated predicted first moon sightings (ref 15). The calculations are based on the rule that the best time and place for making the earliest sightings are when the moon is vertically above the sun at sunset so that their azimuths are equal (ie relative azimuth at sunset=0) and where the apparent altitude of the moon at sunset is 10 degrees. If the sky is clear and the horizon is flat, sighting should be possible just before the sun reaches a geocentric altitude of -5 degrees. The criterion as implemented in MoonCalc is useful for finding the earliest location where the new moon is likely to be sighted. On a global scan, the criterion does *not* show all areas west of the 'earliest point' where the crescent will be seen.

- **South African Astronomical Observatory (SAAO)** [Alt, Rel Azi]

This is a sighting criterion proposed by Drs. John Caldwell and David Laney of the South African Astronomical Observatory (ref 20). The criterion was based on published crescent sightings together with a few local sightings from Signal Hill. The criterion depends on 'topocentric moon altitude (to lower limb) at apparent sunset' and 'difference in azimuth at sunset'. Two lines are drawn on a graph of altitude versus relative azimuth. The sightability of a crescent is 'possible' if above upper line, 'improbable' if between the two line or 'impossible' if below the lower line.

- **Shaukat** [Alt, Crescent width]

This criterion, proposed by Khalid Shaukat and the Committee for Crescent Observation, New York, depends on the 'topocentric altitude of the moon (to the lower limb) at sunset' and the 'calculated crescent width at sunset'. The altitude must be >3.4 degrees at sunset and $(alt/12.7) + (crescent\ width\ in\ arcmin / 1.2) > 1$. The crescent width is calculated in a slightly non-standard way. The criterion has undergone successive refinements based on prospectively collected observation data.

- **Yallop 1997/8** [Rel Alt, Crescent Width]

This criterion was developed from the Indian and Bruin criteria by Bernard Yallop (formerly of the Royal Greenwich Observatory, Cambridge, UK). It takes into account information from 295 published moon (non)sightings compiled by Schaefer and Doggett. The criterion depends on a parameter called 'q' which is derived from the relative geocentric altitude of the moon (ARCV) and topocentric crescent width. This is the default criterion used in MoonCalc. In the original technical note by Yallop (ref 24), q was derived at 'best time' (ie sunset + (4/9)* moonset lag). However, it is not always practical to apply the criterion at 'best time' and so MoonCalc allows the criterion to be applied at sunset or when the sun is at -5 degrees as well as at 'best time'.

The value of q is stratified to give 6 types of predictions:

A: easily visible

B: visible when atmospheric conditions are perfect

C: may need optical aid to find crescent

D: visible with optical aid only

E: not visible even with optical aids

F: outside Danjon limit

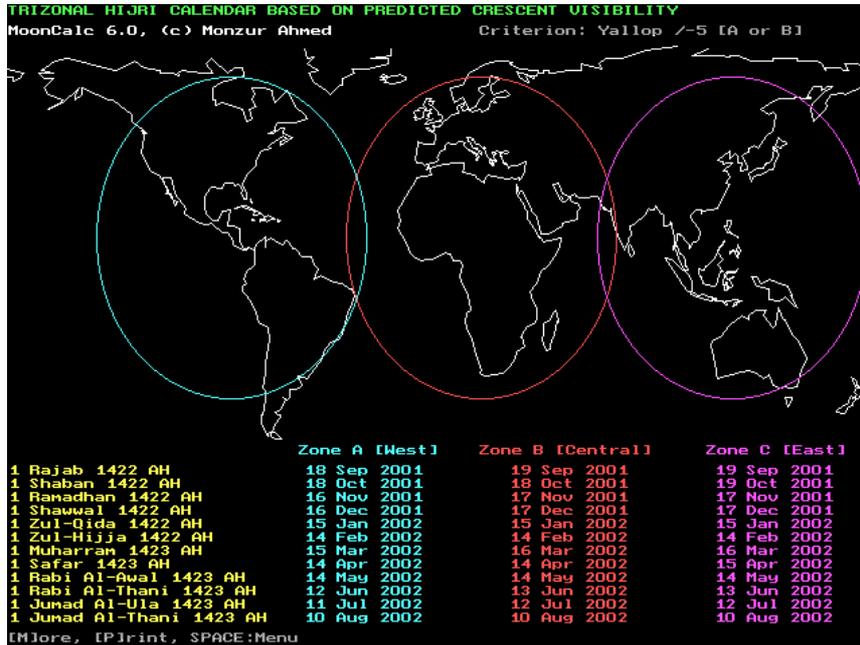
Note that applying the criterion at sunset makes the visibility predictions slightly more pessimistic compared to 'best time', and global visibility zones are west shifted by about 5 degrees of longitude.

- **Schaefer 1988** **Not Yet Implemented**

B.E. Schaefer has developed a complex theoretical sighting criterion based on the idea of Bruin. This criterion apparently takes into account atmospheric haziness, aerosol scattering, Rayleigh scattering, ozone absorption etc (see refs 16,21 and 22). Despite several publications, this criterion has not yet been documented in sufficient detail to implement in MoonCalc.

3.6.3 Trizonal Calendar (map mode)

The calendar generated is the same as in 3.6.2 but the output is in graphic rather than text format. A map of the world is shown with the three zones marked out approximately.

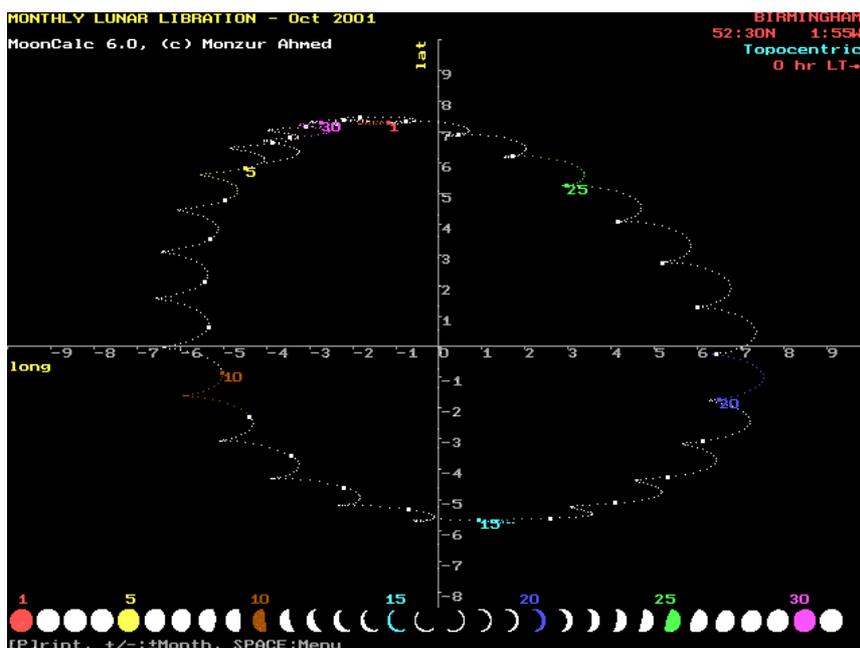


Tri-zonal calendar – the 3 zones are indicated by the coloured ellipses

3.7 Option 7. Libration Graph

Optical libration is the apparent oscillations of the moon due to the variations in the geometric position of the Earth relative to the lunar surface during the course of the orbital motion of the moon. Physical libration is the actual rotational motion of the moon about its mean rotation. Physical libration is much smaller than optical libration and can never be larger than 0.04 degrees in both latitude and longitude.

This option draws a graph showing TOTAL libration in latitude versus TOTAL libration in longitude for a whole month. The user can choose either topocentric or geocentric libration (see 3.10.11).



Libration graph: the dates are labelled on the graph at intervals of 5 days. The square points on the graph represent midnight. At the bottom the screen, the corresponding daily phases of the moon for the month are drawn.

3.10.3 Start and End of Summer Time/Daylight Saving Time

The rules for the start and finish of Summer Time or Daylight Saving Time (DST) vary from country to country. For example, in 1986 the effective periods for DST for various countries were as follows:

COUNTRY	Effective DST period (dates inclusive)		
AUSTRALIA	26 OCT 86	-	28 FEB 86
CANADA	27 APR 86	-	25 OCT 86
FRANCE	30 MAR 86	-	27 SEP 86
IRAQ	01 APR 86	-	30 SEP 86
ITALY	30 MAR 86	-	27 SEP 86
JORDAN	04 APR 86	-	02 OCT 86
SPAIN	30 MAR 86	-	27 SEP 86
SYRIA	16 FEB 86	-	18 OCT 86
TURKEY	30 MAR 86	-	27 SEP 86
USA	27 APR 86	-	25 OCT 86
UK	30 MAR 86	-	25 OCT 86

(NB: for some countries, eg USA, rules may have changed since 1986!)

During DST, one hour (in most countries) is added to the standard time. In many countries there are general rules for the start and end of DST. For example, in the UK, DST (British Summer Time) usually starts on the fourth Sunday of March and ends on the fourth Sunday of October. Similarly, in most areas of the USA, DST starts on the first Sunday of April and ends on the last Sunday of October.

The DST handling of the program has been designed to be flexible enough to cater for most countries of the world. The start/end of DST can be set either as an **absolute** date e.g. 1st May or in a **relative** way e.g. fourth Sunday of March.

Essentially you have to answer 3 questions (following the prompts) to set the start or end of DST:

Q1. The month when DST starts or ends.

Q2. The day on which DST starts or ends.

- for absolute date, choose 'Specific date' for this question.
- for relative date, choose a day name e.g. 'Sunday'

Q3. The position of the day in the month.

- for absolute date, enter the date when you want DST to start/end.
- for relative date, enter the position of the day in the month i.e. first, second, third, fourth or last. For example if you want DST to start on the last Sunday of the chosen month, enter 'last' or if you want DST to start on the fourth Sunday, enter 'fourth'.

Example 1. If you want DST to start on 1st April, the three questions should be answered as follows:

- Q1. 4
- Q2. Specific date
- Q3. 1

Example 2. If you want DST to start on the last Sunday of April, the three questions should be answered as follows:

- Q1. 4
- Q2. Sunday
- Q3. last

The program ships with default start/end of DST valid for the UK i.e. DST starts on the fourth Sunday of March and ends on fourth Sunday of October.

If the Summer Time/DST rules are different for your location then you must alter the rules using this option. If you specify that Summer Time/DST does not apply for your location (when you enter the location into the database) then these rules will be ignored for that location.

3.10.6 Visibility Criterion

The following screen will appear:

WHICH NEW MOON VISIBILITY CRITERION DO YOU WANT TO USE?

0. Babylonian.....Age>24 hrs & Lag>48 mins
1. Ibn Tariq.....[Alt, Lag]
2. Fotheringham.....[Alt, Rel Azi]
3. Maunder.....[Alt, Rel Azi]
4. Indian/Schoch.....[Alt, Rel Azi]
5. Bruin.....[Alt, Crescent Width]
6. Ilyas_A.....[Alt, Elong]
7. Ilyas_B.....[Lag, Latitude]
8. Ilyas_C.....[Alt, Rel Azi]
9. RGO 67.....[Alt, (Rel Azi)]
10. SAAO.....[Alt, Rel Azi]
11. Shaukat.....[Alt, Crescent Width]
12. Yallop 1997/8.....[Rel Alt, Crescent Width]

The current choice is 12

Press ENTER to accept or type in new default (0-12):

Choose the criterion that you wish to use (option 12, Yallop is the default used in the program). See section 3.5.1 for further information on each of these criteria.

If criterion number 9 (RGO 67) is chosen, then 2 further choices are provided:

- Minimum moon altitude at apparent sunset

Generally this should be 10 degrees in line with the RGO recommendations. However the calculated elongation of the world record moon sighting was 8.1 degrees (ref 14). The user is allowed to enter a value in the range 0-25 degrees.

- Maximum relative azimuth at sunset

According to RGO sheet 67 the place/time of earliest moon sighting occurs when the new moon and sun have the same azimuth at sunset ie relative azimuth is zero. When scanning the globe in steps of, say, one or two degrees a relative azimuth of zero is too strict. The program allows the user to 'loosen' the criterion a little by defining the maximum relative azimuth at sunset which can be taken to be zero. (Range allowed 0.001-30 degrees, default: 0.2 degrees, recommended: 0.1-0.5 degrees).

3.10.7 Interval between longitudes

We can define the fineness of the initial grid used for scanning the globe for new moon visibility (option 5 from the main menu). The finer the grid the longer it will take to complete the scan. This option allows you to set the interval in degrees between successive longitudes during the first, coarse scan. Range allowed (1-5 degrees). Use 1 degree if you have a fast computer (486DX or Pentium). Use higher values if you have a slower computer.

3.10.8 Interval between latitudes

This option defines the interval between successive latitudes during the first scan of the globe. The same comments as for 3.10.7 apply.

3.10.9 Lower limit of latitude

To save time you can define the upper and lower limits of latitude for the global scan. Usually a lower limit of -60 degrees (ie 60S) and an upper limit of 60 degrees (60N) are optimal. Range allowed for lower limit -30 to -90 degrees; default -60 degrees.

3.10.10 Upper limit of latitude

See section 3.10.9. Range allowed for upper limit 30-90 degrees; default 60 degrees.

3.10.11 Topocentric or Geocentric

Topocentric = as seen from the observer's place on surface of earth.

Geocentric = as seen from centre of the earth.

For actual moonsighting, it is usual to choose topocentric.

The *displayed* altitudes are always measured to the centre of the moon/sun regardless of topocentric/geocentric setting.

3.10.12 Correction for refraction

Choose "Yes" if you want to compensate for atmospheric refraction.

3.10.13 Apparent or Geometric sunset

Apparent sunset: when the upper limb of the sun is on the horizon taking into account refraction and parallax.

Geometric sunset: when the centre of the sun is on the horizon NOT taking into account refraction or parallax.

Generally this setting should be left on 'Apparent sunset' since this is the "usual" definition of sunset used in civil life.

For expert users: if you want to calculate such values as ARCV (arc of vision), DAZ (difference in azimuth) and ARCL (arc of light) as used in various moon sighting criteria, then set sunset to 'Geometric', Topocentric/Geocentric to 'Geocentric' and Refraction to 'OFF'.... then:

- relative altitude at sunset = ARCV
- relative azimuth at sunset = DAZ
- elongation at sunset = ARCL

3.10.14 Atmospheric Temperature

The value will effect the internal calculation of refraction. Usually this should be set in the range 10-25 degrees Celsius.

3.10.15 Atmospheric Pressure

The value will affect the internal calculation of refraction. Usually this should be set to 1010 millibars.

4. Future Developments

This is the third full release of MoonCalc. However, the program remains in a constant state of development. Please send your suggestions/comments, bug report etc to me either by snail mail or by email (see start of document for addresses).

There are many ideas in the pipeline to enhance MoonCalc. However the constraints of programming in DOS are becoming rather restrictive. I am still working on a Windows version of the program. Maybe, one day I may write a Java version.

At present, my primary concern is to remove any bugs from the data generating engine. As it stands, the program produces reliable data which is compatible with other planetarium type programs and sources.

Generally, the algorithms used in the program are very accurate (based on routines in refs 2 and 4 together with several other sources) and are probably more accurate than is actually needed by most users.

5. Acknowledgements

I should like to thank the many people who helped in the development and testing of this program over the past 8 years. In particular, I should like to thank Shakoor Chughtai (UK) for his helpful comments and extensive error testing in early versions of the program. I am indebted to Omar Afzal (Cornell University, USA) for providing me with certain difficult to locate reference materials. My thanks to Bernard Yallop (formerly of the Royal Greenwich Observatory, UK) for testing MoonCalc and providing friendly advice and stimulating discussions.

I am grateful to the many, many people who made useful comments on the earlier releases of MoonCalc including Rashid Motala, Yusuf Essack, Yaakov Loewinger, Geoff Hitchcox, Robert H. van Gent, Ali Cengia, Bob Cripps of The Eastbourne Astronomical Society (UK), Paul Gabriel, Martin Lewicki, Tariq Muneer, John Taylor, Robert H. Douglass, Jean Meeus, Mohammad Ilyas and Ernst Böck. I should also like to thank my wife, Sayra, for her support and help with digitising the world map.

Finally, a special "hello" to my daughters Zahra and Hanifa aged 4.5 and 2 years respectively when MoonCalc 6.0 was released. They were mainly responsible for the delay in the release of the new version :-)

6. Disclaimer

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7. Glossary

- **Astronomical New Moon**

The moment when the sun, moon and earth are in one plane (in conjunction).

- **Altitude**

The angle up from the horizon. Positive above the horizon, negative below.

- **Azimuth**

The angle around from the north pole measured on the horizon in the sense NESW

- **Bright Limb Angle**

Difficult to explain without a diagram! Imagine a line joining the tips of the two limbs of the bright side of the moon. The BLA is 90 degrees + the anticlockwise angle between the celestial north-south axis and the above - mentioned line.

- **Conjunction**

See astronomical new moon.

- **Declination**

In the equatorial co-ordinate system, the angle measured perpendicular to the equator.

- **Elongation**

The sun-moon elongation is the angular separation of the moon from the sun as observed from a point on earth.

- **Latitude**

The co-ordinate expressing the angle (north positive, south negative) perpendicular to a fundamental plane. On the Earth the geographical longitude is the co-ordinate expressing the angle relative to the equator.

- **Libration**

Optical Libration: apparent oscillations of the moon due to the variations in the geometric position of the Earth relative to the lunar surface during the course of the orbital motion of the moon.

Physical libration: actual rotational motion of the moon about its mean rotation. Physical libration is much smaller than optical libration and can never be larger than 0.04 degrees in both latitude and longitude.

- **Longitude**

The co-ordinate expressing the angle round from a fixed direction measured in a fundamental plane. On the Earth, the geographical longitude is measured with respect to the equator.

- **Magnetic Declination**

The direction of magnetic north relative to true north. If positive, then magnetic north is east of true north. Magnetic Declination varies with time and location.

- **Parallactic angle**

The angle clockwise between the observer's zenith axis and the celestial north-south axis. The parallactic angle will vary with location and time of day. Knowledge of both the parallactic angle and the bright limb angle are needed to determine the orientation of the moon's limbs as observed in the sky.

- **Phase**

The area of the disc (of the moon or a planet) which is illuminated.

- **Positional angle of axis**

Counterclockwise angle between celestial axis and moon's rotational axis.

- **Right Ascension**

in the equatorial co-ordinate system the angle measured around from the point of Aries in the plane of the equator, in the sense SENW.

- **Terrestrial Dynamical Time (TD)**

An uniform time scale for accurate calculations defined by atomic clocks (unlike Greenwich Mean Time and Universal Time which are based on the Earth's rotation). The difference between TD and UT varies with time; currently TD-UT is about 1 minute.

- **Time Zone**

Longitudinal strip on the surface of the earth (approximately 15 degrees of longitude in width) where the zone time is a certain number of hours before or after GMT. This time is adopted as the local civil time by national or international agreement.

8. Abbreviations Used

Alt	Altitude
Azi	Azimuth
BLA	Bright Limb Angle
Dec	Declination
DST	Daylight Saving Time
Elong	Elongation
Geo	Geocentric
GMT	Greenwich Mean Time
hr(s)	Hour(s)
LT	Local Civil Time
Mag Dec	Magnetic Declination
min(s)	Minute(s)
NYI	Not Yet Implemented
RA	Right Ascension
Rel Alt	Relative Altitude
Rel Azi	Relative Azimuth
Semi Diam	Semi-diameter of moon in degrees
sec(s)	Second(s)
TD	Terrestrial Dynamical Time
Topo	Topocentric
TZ	Time Zone
Width	Crescent width in minutes
UT	Universal Time
*	1 hour added for DST/Summer Time

9. General Moon Information

- Distance from Earth:
 - centre to centre: mean: 384,400km
 - closest (perigee): 356,410km
 - furthest (apogee): 406,697km
 - surface to surface: mean: 376,284km
 - closest (perigee): 348,294km
 - furthest (apogee): 398,581km
- Revolution period: 27.321661 days
- Axial rotation period: 27.321661 days
- Synodic period: 29d 12h 44m 2.9s
- Mean orbital velocity: 3680km/h
- Axial inclination of equator, referred to ecliptic: 1d 32m
- Orbital inclination: 5d 09m
- Orbital eccentricity: 0.0549
- Diameter: 3475.6km
- Apparent diameter seen from Earth:
 - max 33m 31s
 - min 29m 22s
 - mean 31m 5s
- Reciprocal mass, Earth = 1: 81.3
- Mass = 7.35×10^{25} g
- Mass, Earth = 1: 0.0123
- Volume, Earth =1: 0.0203
- Escape Velocity = 2.38 km/s
- Surface Gravity, Earth = 1:0.1653
- Albedo: 0.07
- Mean magnitude at full moon: -12.7

10. Frequently Asked Questions

Q. How do I capture and save a MoonCalc map or other MoonCalc graphics screen?

A. The easiest way is to run MoonCalc under Windows. When MoonCalc has produced a graphics screen, press the "Print Screen" key (usually next to "Scroll Lock"). This captures the image to the clipboard. Next run a graphics program such as Paint Shop Pro. "Paste in" the captured image. Then reduce the image to 16 colours and save as a gif file or bmp file. The above capture trick works for resolutions upto 640x480 (i.e. not for 600x800 and 1024x768 modes).

Q. How do I print out MoonCalc data?

A. To printout the tables of data: on some systems running in pure DOS mode, pressing the "Print Screen" button should result in a printout. If the program is running in a DOS box from Windows, pressing "Print Screen" will make a copy of the table to the clipboard. Open Windows Notepad and "Paste in" the clipboard using Control & V. Make sure you are using a mono-spaced font such as Courier. You can now printout the Notepad screen. To printout MoonCalc graphics screen: on a pure DOS system pressing 'P' after a graphics screen has been printed out gives the option of printing out to a Epson or HP Laserjet printer. In Windows mode, you can capture the graphics as described in the previous question and printout using a program such as Paint Shop Pro (works for up to 640x480 resolution).

Q. Why are the sunset, sunrise, moonset and moonrise times out by an hour?

A. Either the time zone is entered incorrectly or daylight saving time rules have to be changed. The Time Zone that MoonCalc suggests are only approximate.

Q. Why doesn't MoonCalc run properly under Windows 95/8?

A. Make sure that the "start in:" or "Working directory" property of the desktop shortcut to the moonc60.exe file points to the directory containing the MoonCalc files.

Q. Why does MoonCalc crash with "runtime error 200" on my fast PC?

A. This should not occur with MoonCalc 5 and higher although it used to occur with older version of the program. Basically, the error was due to a bug in the compiler used to write MoonCalc which caused an overflow when the program was ran on a fast Pentium computer. I have tested version 6.0 on a variety of fast PCs without problems.

Q. Why does MoonCalc crash with "runtime error 207" on my PC?

A. This is a erratic problem which seems to be related to the magnetic declination routine. I have not got to the bottom of it. Often running the program again sorts things out. If the error keeps occurring, try moving the magmodel.dat file out of the working directory - this bypasses the magnetic declination routines.

11. Conclusions

MoonCalc was developed over a period of 8 years and continues to be in a state of constant development. The program has given me a lot of pleasure to write and I hope very much that you enjoy using MoonCalc and find it useful.

Users of MoonCalc are encouraged to test the various functions of the program and compare the data produced by the program with actual sightings of the moon, particularly sightings of the crescent moon. All suggestions and comments which may improve the program are welcomed.

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<http://www.starlight.demon.co.uk/mooncalc>
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10th October 2001

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