## THE 4D/5D CONNECTION BLACK HOLES and HIGHER-DERIVATIVE COUPLINGS

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## Motivation:

What is the relation between:

## microscopic/statistical entropy

macroscopic/field-theoretic entropy
准 microstate counting $\longrightarrow$ entropy $\quad S_{\text {micro }}=\ln d(q, p)$
次 supergravity: Noether surface charge wald, 1993 first law of black hole mechanics (BH thermodynamics)

A reliable comparison requires a (super)symmetric setting!
Ideal testing ground: supergravities with 8 supercharges

- $D=4$ space-time dimensions with $N=2$ supersymmetry
- $D=5$ space-time dimensions with $N=1$ supersymmetry


## To clarify previous results on 4D and 5D black hole solutions, to further the understanding between them, and to obtain new results.

- Higher-derivative terms (4D and 5D) Bergshoeff, de Roo, aw, 1981 Chern-Simons terms (5D) Cardoso, dW, Mohaupt, 1998 Hanaki, Ohashi, Tachikawa, 2007
dW, Katmadas, van Zalk, 2009, 2010
-5D black holes (and black rings)
- the 4D/5D connection
- more recent developments

Castro, Davis, Kraus, Larsen, 2007, 2008 dW, Katmadas, 2009

Gaiotto, Strominger, Yin, 2005
Behrndt, Cardoso, Mahapatra, 2005

Sen, 2011
Banerjee, dW, Katmadas, 2011

## Methodology: superconformal multiplet calculus

- off-shell irreducible supermultiplets in superconformal gravity background
- extra superconformal gauge invariances
- gauge equivalence (compensating supermultiplets)



## BPS black holes in four space-time dimensions

$N=2$ supergravity: vector multiplet sector (Wilsonian effective action) vector multiplets contain scalars $X^{I}$
projectively defined: $X^{I} \longrightarrow Y^{I} \quad$ (residual scale invariance)
Lagrangian encoded in a holomorphic homogeneous function

$$
F(\lambda Y)=\lambda^{2} F(Y)
$$

with near-horizon geometry: $\mathrm{AdS}_{2} \times S^{2}$

## Subleading corrections to Bekenstein-Hawking area law:

 extend with one 'extra' complex field, originating from pure supergravity

$$
\text { homogeneity: } F\left(\lambda Y, \lambda^{2} \Upsilon\right)=\lambda^{2} F(Y, \Upsilon)
$$

$\Upsilon$-dependence leads to terms $\propto\left(R_{\mu \nu \rho}{ }^{\sigma}\right)^{2}$ in effective action BPS: supersymmetry at the horizon

$$
\begin{aligned}
Y^{I}-\bar{Y}^{I} & =\mathrm{i} p^{I} & & \text { magnetic charges } \\
F_{I}-\bar{F}_{I} & =\mathrm{i} q_{I} & & \text { electric charges }
\end{aligned}
$$

Ferrara, Kallosh, Strominger, 1996 Cardoso, dW, Käppeli, Mohaupt, 2000
covariant under dualities! refers to the full function $F$ !

Furthermore $\Upsilon=-64$ leads to subleading corrections

The Wilsonian effective action is not sufficient to realize all dualities. When integrating out the massless modes, so as to obtain the 1PI action, one encounters non-holomorphic corrections. For the BPS near-horizon region one has a natural infrared cut-off provided by the $S^{2}$ radius.

In Minkowski space-time the integration over massless modes is problematic!

It is possible to determine the short-distance corrections that depend logarithmically on the $S^{2}$ radius. Sen, 2011

Surprisingly enough these are consistent with the terms found from a study of degeneracies of BPS states of D-branes on compact Calabi-Yau manifolds.

Denef, Moore, 2007
There is another issue related to the coupling to Gauss-Bonnet terms, which has been used to obtain the entropy for heterotic black holes. However, their supersymmetric form is not known.

Sen, 2005

## BPS black holes and rings in five space-time dimensions

two different supersymmetric horizon topologies !
$\begin{array}{lll}\lessgtr & S^{3} \quad \text { (SPINNING) BLACK HOLE } & \text { Breckenridge, Myers, Peet, Vafa, } 1996 \\ \lessgtr & S^{1} \times S^{2} \text { BLACK RING } & \text { Elvang, Emparan, Mateos, Reall, } 2004\end{array}$
with near-horizon geometry: $\mathrm{AdS}_{2} \times S^{3}$ or $\mathrm{AdS}_{3} \times S^{2}$
(this result does not depend on the details of the Lagrangian)
$5 D$ vector supermultiplets contain $\left\{\begin{array}{lc}\text { scalar fields: } & \sigma^{I} \\ \text { vector fields: } & W_{\mu}{ }^{I}\end{array}\right.$ abelian field strengths $F_{\mu \nu}{ }^{I}=\partial_{\mu} W_{\nu}{ }^{I}-\partial_{\nu} W_{\mu}{ }^{I}$
supergravity (Weyl) multiplet contains
(auxiliary) tensor $T_{\mu \nu}$$\rightarrow T_{a b}\left\{\begin{array}{l}T_{01} \\ T_{23}\end{array}\right.$

$$
v^{2} \equiv\left(T_{01}\right)^{2}+\left(T_{23}\right)^{2}
$$

## supersymmetry + partial gauge choice

$$
\sigma^{I}=\text { constant } \quad \text { (remain subject to residual (constant) scale transformations!) }
$$

$T_{\mu \nu}$ conformal Killing-Yano tensor

$$
\mathcal{D}_{\rho} T_{\mu \nu}=\frac{1}{2} g_{\rho[\mu} \xi_{\nu]} \quad \xi^{\mu} T_{\mu \nu}=0
$$

Killing vector associated with the fifth dimension $\psi$

$$
\xi^{\mu} \propto e^{-1} \varepsilon^{\mu \nu \rho \sigma \tau} T_{\nu \rho} T_{\sigma \tau}
$$

$$
\begin{aligned}
d s^{2} & =\frac{1}{16 v^{2}}\left(-r^{2} \mathrm{~d} t^{2}+\frac{\mathrm{d} r^{2}}{r^{2}}+\mathrm{d} \theta^{2}+\sin ^{2} \theta \mathrm{~d} \varphi^{2}\right)+\mathrm{e}^{2 g}(\mathrm{~d} \psi+\sigma)^{2} \\
\sigma & =-\frac{1}{4 v^{2}} \mathrm{e}^{-g}\left(T_{23} r \mathrm{~d} t-T_{01} \cos \theta \mathrm{~d} \varphi\right)
\end{aligned}
$$

## Two distinct cases:

- $T_{01} \neq 0 \quad$ SPINNING BLACK HOLE

$$
\text { angular momentum } \propto \frac{T_{23}}{T_{01}}
$$

- $T_{01}=0 \quad$ BLACK RING

Additional horizon condition and 'magnetic' charges

$$
\begin{aligned}
& F_{\mu \nu}^{I}=4 \sigma^{I} T_{\mu \nu} \quad \mathcal{Q}_{\mu \nu}=\partial_{\mu} \sigma_{\nu}-\partial_{\nu} \sigma_{\mu} \\
& F_{\theta \varphi} \longrightarrow \quad p^{I}=\frac{\sigma^{I}}{4 v^{2}} T_{23} \\
& \mathcal{Q}_{\theta \varphi} \longrightarrow \quad p^{0}=\frac{\mathrm{e}^{-g}}{4 v^{2}} T_{01}
\end{aligned} \quad \begin{aligned}
& \text { attractor equations }
\end{aligned}
$$

## Scale invariance (residue of conformal invariance)

$\sigma^{I}, T_{a b}, v, \mathrm{e}^{-g} \quad$ scale uniformly
the metric is scale dependent

Action in 5 space-time dimensions consists of two cubic invariants, each containing a Chern-Simons term:

$$
\begin{aligned}
& \mathcal{L} \propto C_{I J K} \varepsilon^{\mu \nu \rho \sigma \tau} W_{\mu}{ }^{I} F_{\nu \rho}{ }^{J} F_{\sigma \tau}{ }^{K} \\
& \mathcal{L} \propto c_{I} \varepsilon^{\mu \nu \rho \sigma \tau} W_{\mu}{ }^{I} R_{\nu \rho}{ }^{a b} R_{\sigma \tau} a b
\end{aligned}
$$

Hanaki, Ohashi, Tachikawa, 2006
dW,Katmadas, 2009
The Chern-Simons terms cause non-trivial complications in the determination of entropy, electric charges and angular momenta!

## An example: 5D electromagnetism with CS term

$\mathcal{L}^{\text {total }}=\mathcal{L}^{\text {inv }}\left(F_{\mu \nu}, \nabla_{\rho} F_{\mu \nu}, \psi, \nabla_{\mu} \psi\right)+\varepsilon^{\mu \nu \rho \sigma \tau} A_{\mu} F_{\nu \rho} F_{\sigma \tau}$
The Noether potential associated with the abelian gauge symmetry takes the form

$$
\left.\mathcal{Q}_{\text {gauge }}^{\mu \nu}(\phi, \xi)=2 \mathcal{L}_{F}^{\mu \nu} \xi-2 \nabla_{\rho} \mathcal{L}_{F}^{\rho, \mu \nu} \xi+6 e^{-1} \varepsilon^{\mu \nu \rho \rho \sigma \tau} \uparrow\right\} A_{\rho} F_{\sigma \tau}
$$

where $\delta \mathcal{L}^{\text {inv }}=\mathcal{L}_{F}^{\mu \nu} \delta F_{\mu \nu}+\mathcal{L}_{F}^{\rho, \mu \nu} \delta\left(\nabla_{\rho} F_{\mu \nu}\right)+\mathcal{L}_{\psi} \delta \psi+\mathcal{L}_{\psi}^{\mu} \delta\left(\nabla_{\mu} \psi\right)$

$$
\partial_{\nu} \mathcal{Q}_{\text {gauge }}^{\mu \nu}=J_{\text {Noether }}^{\mu}=0
$$

* 

$\mathcal{Q}_{\text {gauge }}$ is a closed (d-2)-form for symmetric configurations!
Electric charge is defined as

$$
q=\int_{\Sigma_{\text {hor }}} \varepsilon_{\mu \nu} \mathcal{Q}_{\substack{\text { biauge } \\ \text { bi-normal }}}^{\mu \nu}(\phi, \xi)
$$

(This definition coincides with the definition based on the field equations! )
The electric charge now contains the integral over a 3-cycle of the CS term!

This poses no difficulty for black holes for which the gauge fields are globally defined.
However, the mixed CS term leads to the integral over a gravitational CS term, which is problematic.

For black rings the gauge fields are not globally defined. Both CS terms are therefore problematic.

## Entropy and angular momentum

Entropy (based on first law of black hole mechanics
$\mathcal{S}_{\text {macro }}=-\left.\pi \int_{\Sigma_{\text {hor }}} \varepsilon_{\mu \nu} \mathcal{Q}^{\mu \nu}(\xi)\right|_{\nabla_{[\mu} \xi_{\nu]}=\varepsilon_{\mu \nu} ; \xi^{\mu}=0}$ Wald, 1993
$\xi^{\mu} \partial_{\mu}=\partial / \partial t$ timelike Killing vector bi-normal
Angular momenta
$J(\xi)=\int_{\Sigma_{\text {hor }}} \varepsilon_{\mu \nu} \mathcal{Q}^{\mu \nu}(\xi)$
$\xi^{\mu}$ periodic Killing vector

## Evaluating the CS terms for black rings

The correct evaluation of the CS term for the ring geometry yield

$$
Q_{I}^{\mathrm{CS}} \propto \oint_{\Sigma} C_{I J K} W^{J} \wedge F^{K} \propto C_{I J K} a^{J} p^{K}
$$

Hence, integer shifts of the Wilson line moduli induce a shift in the integrated CS term
For concentric rings, one finds

$$
Q_{I}^{\mathrm{CS}}-6 C_{I J K} P^{J} P^{K}=-12 C_{I J K} \sum_{i}\left(a^{J}+\frac{1}{2} p^{J}\right)_{i} p^{K}{ }_{i}
$$

$$
\text { with } P^{I}=\sum_{i} p_{i}^{J}
$$

The integrated CS terms are not additive! Hanaki, Ohashi, Tachikawa, 2007
Confirmed by explicit results for global solutions.
Gauntlett, Gutowski, 2004
Additive charges take the following form
(upon solving the Wilson line moduli in terms of the charges)

$$
q_{I}-6 C_{I J K} p^{J} p^{K}
$$

## $5 D$ black ring versus $4 D$ black hole:

$$
\begin{gathered}
\mathcal{S}_{\text {macro }}^{\mathrm{BR}}=\frac{4 \pi}{\phi^{0}}\left[C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{4} c_{I} p^{I}\right] \quad D=5 \\
q_{I}=-12 C_{I J K} p^{J} a^{K} \\
J_{\varphi}=p^{I}\left(q_{I}-\frac{1}{6} C_{I J} p^{J}\right) \\
J_{\psi}-J_{\varphi}-\frac{1}{24} C^{I J}\left(q_{I}-6 C_{I K} p^{K}\right)\left(q_{J}-6 C_{J L} p^{L}\right)=\frac{2}{\phi^{0^{2}}}\left[C_{I J K} p^{I} p^{J} p^{K}+\frac{1}{4} c_{I} p^{I}\right] \\
\text { up to calibration } \\
\mathcal{S}_{4 \mathrm{D}}^{\mathrm{BH}}=-\frac{2 \pi}{\phi^{0}}\left[D_{I J K} p^{I} p^{J} p^{K}+256 d_{I} p^{I}\right] \\
q_{I}^{4 \mathrm{D}}=\frac{6}{\phi^{0}} D_{I} \int_{J K} p^{J} \phi^{K} \\
\hat{q}_{0}{ }^{4 D} \equiv q_{0}^{4 D}+\frac{1}{12} D^{I J} q_{I} q_{J}=\frac{1}{\phi^{0^{2}}}\left[D_{I J K} p^{I} p^{J} p^{K}+256 d_{I} p^{I}\right]
\end{gathered}
$$

## COMMENTS:

Confirmation from near-horizon analysis in the presence of higher-derivative couplings. Partial results were already known (but somewhat disputed at the time).

The Wilson line moduli are defined up to integers. This implies that the electric charges and angular momenta are shifted under the large gauge transformations (spectral flow) induced by these integer shifts. Indeed, under

$$
a^{I} \rightarrow a^{I}+k^{I}
$$

one finds,

$$
\begin{aligned}
q_{I} & \rightarrow q_{I}-12 C_{I J K} p^{J} k^{K} \\
J_{\varphi} & \rightarrow J_{\varphi}-12 C_{I J K} p^{I} p^{J} k^{K} \\
J_{\psi} & \rightarrow J_{\psi}-q_{I} k^{I}-6 C_{I J K} p^{I} p^{J} k^{K}+6 C_{I J K} p^{I} k^{J} k^{K}
\end{aligned}
$$

These transformations are in agreement with the corresponding $4 D$ black holes where the above transformations correspond to a duality invariance!

4D/5D connection: difference resides in the contributions from the Chern-Simons term!

## The spinning black hole

$$
\begin{aligned}
\mathcal{S}_{\text {macro }}^{\mathrm{BH}} & =\frac{\pi \mathrm{e}^{g}}{4 v^{2}}\left[C_{I J K} \sigma^{I} \sigma^{J} \sigma^{K}+4 c_{I} \sigma^{I} T_{23}^{2}\right] \\
q_{I} & =\frac{6 \mathrm{e}^{g}}{4 T_{01}}\left[C_{I J K} \sigma^{J} \sigma^{K}-c_{I} T_{01}^{2}\right] \\
p^{0} & =\frac{\mathrm{e}^{-g}}{4 v^{2}} T_{01} \\
J_{\psi} & =\frac{T_{23} \mathrm{e}^{2 g}}{T_{01}^{2}}\left[C_{I J K} \sigma^{I} \sigma^{J} \sigma^{K}-4 c_{I} \sigma^{I} T_{01}^{2}\right]
\end{aligned}
$$

choose scale invariant variables

$$
\begin{aligned}
\phi^{I} & =\frac{\sigma^{I}}{4 T_{01}} \\
\phi^{0} & =\frac{e^{-g} T_{23}}{4 v^{2}}=\frac{p^{0} T_{23}}{T_{01}}
\end{aligned}
$$

$$
\mathcal{S}_{\text {macro }}^{\mathrm{BH}}=\frac{4 \pi p^{0}}{\left(\phi^{0^{2}}+p^{0^{2}}\right)^{2}}\left[p^{0^{2}} C_{I J K} \phi^{I} \phi^{J} \phi^{K}+\frac{1}{4} c_{I} \phi^{I} \phi^{0^{2}}\right]
$$

$$
q_{I}=\frac{6 p^{0}}{\phi^{0^{2}}+p^{0^{2}}}\left[C_{I J K} \phi^{J} \phi^{K}-\frac{1}{16} c_{I}\right] \quad D=5
$$

$$
J_{\psi}=\frac{4 \phi^{0} p^{0}}{\left(\phi^{0^{2}}+p^{0^{2}}\right)^{2}}\left[C_{I J K} \phi^{I} \phi^{J} \phi^{K}-\frac{1}{4} c_{I} \phi^{I}\right]
$$

up to calibration relative factor $\frac{4}{3}$

$$
\begin{aligned}
\mathcal{S}_{4 D}^{\mathrm{BH}} & =\frac{2 \pi p^{0}}{\left(\phi^{0^{2}}+p^{0^{2}}\right)^{2}}\left[p^{0^{2}} D_{I J K} \phi^{I} \phi^{J} \phi^{K}+256 d_{I} \phi^{I} \phi^{0^{2}}\right] \\
q_{I}^{4 D} & =-\frac{3 p^{0}}{\phi^{0^{2}}+p^{0^{2}}}\left[D_{I J K} \phi^{J} \phi^{K}-\frac{256}{3} d_{I}\right] \quad D=4 \\
q_{0}^{4 D} & =\frac{2 \phi^{0} p^{0}}{\left(\phi^{0^{2}}+p^{0^{2}}\right)^{2}}\left[D_{I J K} \phi^{I} \phi^{J} \phi^{K}-256 d_{I} \phi^{I}\right]
\end{aligned}
$$

## Comments:

Agrees with microstate counting for $J_{\psi}=0$
Vafa, 1997
Huang, Klemm, Mariño, Tavanfar, 2007
Two important differences with the literature!
Castro, Davis, Kraus, Larsen, 2007,2008
The expression for $J_{\psi}$ is rather different! We have $J_{\psi}=q_{0}$.
The electric charges receives different corrections from the higher-order derivative couplings:

$$
3 c^{I} \Leftrightarrow\left[3-\frac{\phi^{0^{2}}}{p^{0^{2}}}\right] c_{I} \quad \text { identical when } J_{\psi}=0
$$

It turns out that the discrepancy between 4D/5D results resides in the mixed Chern Simons term.

$$
\text { Banerjee, dW, Katmadas, } 2011
$$

To exclude the possibility that contributions from other terms could be present one has to make a detailed comparison of $4 D$ and $5 D$ supergravity.

Inl| the off-shell 4D/5D connection

The off-shell 4D/5D connection also enables one to understand why the field equations of the 5D and the 4D theory are different (in the bulk), as has been observed in the literature.

Upon dimensional reduction the invariant Lagrangians with at most two derivatives lead to the corresponding 4D Lagrangians. There will be no qualitative differences.

However, the higher-order derivative term in 5D leads upon reduction to three separate 4D invariant Lagrangians.

One of these Lagrangians is known to give no contribution to the entropy and the electric charges of 4D BPS black holes, because of a non-renormalization theorem.
dW, Katmadas, van Zalk, 2010

## Off-shell dimensional reduction

Weyl multiplet, vector multiplet, hypermultiplet
First Kaluza-Klein ansatz with gauge choices:

$$
e_{M}^{A}=\left(\begin{array}{cc}
e_{\mu}^{a} & B_{\mu} \phi^{-1} \\
0 & \phi^{-1}
\end{array}\right), \quad e_{A}^{M}=\left(\begin{array}{cc}
e_{a}^{\mu} & -e_{a}^{\nu} B_{\nu} \\
0 & \phi
\end{array}\right), \quad b_{M}=\binom{b_{\mu}}{0}
$$

There are compensating transformations to preserve this form.
Upon the reduction the 5D Weyl multiplet decomposes into the $4 D$ Weyl multiplet and a Kaluza-Klein vector multiplet.

Continue with the other fields and then consider the action of the five-dimensional supersymmetry transformations.

In this case, there is no conflict between conformal invariance and dimensional reduction.

The supersymmetry transformations turn out to satisfy a uniform decomposition:
$\left.\delta_{\mathrm{Q}}(\epsilon)\right|_{5 D} ^{\text {reduced }} \Phi=\left.\delta_{\mathrm{Q}}(\epsilon)\right|_{4 D} \Phi+\left.\delta_{\mathrm{S}}(\tilde{\eta})\right|_{4 D} \Phi+\left.\delta_{\mathrm{SU}(2)}(\tilde{\Lambda})\right|_{4 D} \Phi+\left.\delta_{\mathrm{U}(1)}\left(\tilde{\Lambda}^{0}\right)\right|_{4 D} \Phi$.

The Kaluza-Klein vector multiplet has a real scalar field, rather than a complex one. Moreover, the R-symmetry remains restricted to $\mathrm{SU}(2)$ rather than extending to $\mathrm{SU}(2) \times \mathrm{U}(1)$ !

Explanation: It turns out the the result takes the form of a gauge-fixed version with respect to $\mathrm{U}(1)$ !
compensating transformation

## Weyl multiplet 5D $\Rightarrow$ 4D <br> (reducible)

$$
\begin{aligned}
& \phi=2\left|X^{0}\right| \quad \begin{array}{l}
\text { KK vector multitiplet } \\
\text { missing pseudloscalar }
\end{array} \\
& B_{\mu}=W_{\mu}^{0} \quad \\
& V_{M i}^{j}=\left\{\begin{array}{l}
V_{\mu i}{ }^{j}=\mathcal{V}_{\mu}^{j} i_{i}-\frac{1}{4} \varepsilon_{i k} Y^{k j 0}\left|X^{0}\right|^{2} W_{\mu}^{0} \\
V_{5 i}{ }^{j}=-\frac{1}{4} \varepsilon_{i k} Y^{k j 0}\left|X^{0}\right|^{2}
\end{array}\right. \\
& T_{a 5}=\frac{1}{12} \mathrm{i}_{a}{ }^{\mu}\left(\frac{\mathcal{D}_{\mu} X^{0}}{X^{0}}-\frac{\mathcal{D}_{\mu} \bar{X}^{0}}{\bar{X}^{0}}\right) \\
& T_{a b}=-\frac{\mathrm{i}}{24\left|X^{0}\right|}\left(\varepsilon_{i} T_{a b}^{i j} \bar{X}^{0}-F_{a b}^{-}(B)\right)+\text { h.c. }
\end{aligned}
$$

## vector multiplet $5 D=4 D$

$$
\begin{aligned}
& \sigma=-\mathrm{i}\left|X^{0}\right|(t-\bar{t}) \\
& W_{M}=\left\{\begin{array}{l}
W_{\mu}=W_{\mu}-\frac{1}{2}(t+\bar{t}) W_{\mu}^{0} \\
W_{5}=-\frac{1}{2}(t+\bar{t})
\end{array}\right. \\
& Y^{i j}=-\frac{1}{2} Y^{i j}+\frac{1}{4}(t+\bar{t}) Y^{i j 0} \\
& t=\frac{X}{X^{0}}
\end{aligned}
$$

## 5D vector multiplet action

$$
\begin{aligned}
8 \pi^{2} \mathcal{L}_{\mathrm{vvv}}= & 3 E C_{A B C} \sigma^{A}\left[\frac{1}{2} \mathcal{D}_{M} \sigma^{B} \mathcal{D}^{M} \sigma^{C}+\frac{1}{4} F_{M N}{ }^{B} F^{M N C}-Y_{i j}{ }^{B} Y^{i j C}-3 \sigma^{B} F_{M N}{ }^{C} T^{M N}\right] \\
& -\frac{1}{8} \mathrm{i} C_{A B C} \varepsilon^{M N P P Q R} W_{M}{ }^{A} F_{N P}{ }^{B} F_{Q R}{ }^{C} \\
& -E C_{A B C} \sigma^{A} \sigma^{B} \sigma^{C}\left[\frac{1}{8} R-4 D-\frac{39}{2} T^{A B} T_{A B}\right]
\end{aligned}
$$

## 5D higher-derivative action

$$
\begin{aligned}
8 \pi^{2} \mathcal{L}_{\mathrm{vww}}= & \frac{1}{4} E c_{A} Y_{i j}^{A} T^{C D} R_{C D k}{ }^{j}(V) \varepsilon^{k i} \\
& +E c_{A} \sigma^{A}\left[\frac{1}{64} R_{C D}{ }^{E F}(M) R_{E F}{ }^{C D}(M)+\frac{1}{96} R_{M N j}{ }^{i}(V) R^{M N{ }_{i}}{ }^{j}(V)\right] \\
& -\frac{1}{128} \mathrm{i} \varepsilon^{M N P Q R} c_{A} W_{M}^{A}\left[R_{N P}{ }^{C D}(M) R_{Q R C D}(M)+\frac{1}{3} R_{N P j}{ }^{i}(V) R_{Q R i}{ }^{j}(V)\right] \\
& +\frac{3}{16} E c_{A}\left(10 \sigma^{A} T_{B C}-F_{B C}{ }^{A}\right) R(M)_{D E}{ }^{B C} T^{D E}+\cdots
\end{aligned}
$$

## yields the following 4D actions:

$$
F\left(X,\left(T_{a b}{ }^{i j} \varepsilon_{i j}\right)^{2}\right)=-\frac{1}{2} \frac{C_{A B C} X^{A} X^{B} X^{C}}{X^{0}}-\frac{1}{2048} \frac{c_{A} X^{A}}{X^{0}}\left(T_{a b}{ }^{i j} \varepsilon_{i j}\right)^{2}
$$

corresponding to the chiral superspace density invariant

$$
\mathcal{H}(X, \bar{X})=\frac{1}{384} \mathrm{i} c_{A}\left(\frac{X^{A}}{X^{0}} \ln \bar{X}^{0}-\frac{\bar{X}^{A}}{\bar{X}^{0}} \ln X^{0}\right)
$$

corresponding to a full superspace density invariant, described in terms of a Kähler potential. This class of actions does not contribute to the electric charges and the entropy of BPS black holes.

## Supersymmetric 4D Lagrangian arising from dimensional reduction of five-dimensional supergravity:

$$
\begin{aligned}
& e^{-1} \mathcal{L}= \\
& \mathcal{H}_{I J \bar{K} \bar{L}}\left[\frac{1}{4}\left(F_{a b}^{-I} F^{-a b J}-\frac{1}{2} Y_{i j}{ }^{I} Y^{i j J}\right)\left(F_{a b}^{+K} F^{+a b L}-\frac{1}{2} Y^{K i j} Y^{L}{ }_{i j}\right)\right. \\
& \left.+4 \mathcal{D}_{a} X^{I} \mathcal{D}_{b} \bar{X}^{K}\left(\mathcal{D}^{a} X^{J} \mathcal{D}^{b} \bar{X}^{L}+2 F^{-a c J} F^{+b_{c}}{ }_{c}{ }^{2}-\frac{1}{4} \eta^{a b} Y_{i j}^{J} Y^{L i j}\right)\right] \\
& +\left\{\mathcal { H } _ { I J \overline { K } } \left[4 \mathcal{D}_{a} X^{I} \mathcal{D}^{a} X^{J} \mathcal{D}^{2} \bar{X}^{K}-\left(F^{-a b I} F_{a b}^{-J}-\frac{1}{2} Y_{i j}^{I} Y^{J i j}\right)\left(\square_{\mathrm{c}} X^{K}+\frac{1}{8} F_{a b}^{-K} T^{a b i j} \varepsilon_{i j}\right)\right.\right. \\
& \left.\left.+8 \mathcal{D}^{a} X^{I} F_{a b}^{-J}\left(D_{c} F^{+c b K}-\frac{1}{2} \mathcal{D}_{c} \bar{X}^{K} T^{i j c b} \varepsilon_{i j}\right)-\mathcal{D}_{a} X^{I} Y_{i j}^{J} \mathcal{D}^{a} Y^{K i j}\right]+ \text { h.c. }\right\} \\
& +\mathcal{H}_{I \bar{K}}\left[4\left(\square_{\mathrm{c}} \bar{X}^{I}+\frac{1}{8} F_{a b}^{+I} T^{a b}{ }_{i j} \varepsilon^{i j}\right)\left(\square_{\mathrm{c}} X^{K}+\frac{1}{8} F_{a b}^{-K} T^{a b i j} \varepsilon_{i j}\right)+4 \mathcal{D}^{2} X^{I} \mathcal{D}^{2} \bar{X}^{K}\right. \\
& +8 \mathcal{D}_{a} F^{-a b I} \mathcal{D}_{c} F^{+c}{ }_{b}{ }^{K}-\mathcal{D}_{a} Y_{i j}{ }^{I} \mathcal{D}^{a} Y^{K i j}+\frac{1}{4} T_{a b}{ }^{i j} T_{c d i j} F^{-a b I} F^{+c d} K \\
& +\left(\frac{1}{6} R(\omega, e)+2 D\right) Y_{i j}{ }^{I} Y^{i j K}+4 F^{-a c I} F^{+}{ }_{b c}{ }^{K} R(\omega, e)_{a}{ }^{b} \\
& -\left[\varepsilon^{i k} Y_{i j}{ }^{I} F^{+a b K} R(\mathcal{V})_{a b}{ }^{j}{ }_{k}+\text { h.c. }\right] \\
& -\left[\mathcal{D}_{c} \bar{X}^{K}\left(D^{c} T_{a b}{ }^{i j} F^{-I a b}+4 T^{i j c b} D^{a} F_{a b}^{-I}\right) \varepsilon_{i j}+\text { h.c. }\right] \\
& \left.+8\left(R^{\mu \nu}(\omega, e)-\frac{1}{3} g^{\mu \nu} R(\omega, e)+\frac{1}{4} T^{\mu}{ }_{b}{ }^{i j} T^{\nu b}{ }_{i j}+\mathrm{i} R(A)^{\mu \nu}-g^{\mu \nu} D\right) \mathcal{D}_{\mu} X^{I} \mathcal{D}_{\nu} \bar{X}^{K}\right]
\end{aligned}
$$

## Surprise: there is one more higher-derivative coupling emerging from 5D !

Some characteristic terms:

$$
\begin{aligned}
8 \pi^{2} \mathcal{L}_{\mathrm{VWw}} \rightarrow & -\frac{1}{384} \mathrm{i} c_{A} t^{A}\left[\frac{2}{3} \mathcal{R}_{a b} \mathcal{R}^{a b}+R(\mathcal{V})_{a b j}^{+i} R(\mathcal{V})^{+a b j}{ }_{i}\right] \\
& -\frac{1}{768} \mathrm{i} c_{A}\left(t^{A}-\bar{t}^{A}\right)\left(X^{0}\right)^{-1} \varepsilon_{i j} T^{c d i j} R(M)^{a b}{ }_{c d} F_{a b}^{-0} \\
& + \text { h.c. }
\end{aligned}
$$

where $\mathcal{R}_{\mu \nu}$ is the Ricci tensor

Its structure is not yet known. Neither is it known whether this coupling is subject to some non-renormalization theorem. It will probably have some bearing on the supersymmetric Gauss-Bonnet invariant.

## Conclusions/comments

Off-shell dimensional reduction enables one to get a precise comparison, also in the presence of higher-order invariants, between complicated Lagrangians in 4 and 5 dimensions.

The known higher-derivative invariant coupling in 5D gives rise to three independent higher-derivative couplings in 4D. The indication is that only one of them contributes to electric charges and entropy of BPS black holes. One is not yet explicitly known, and seems to be related to a supersymmetric coupling to the Gauss-Bonnet term.

The difference in the near-horizon behavior between 4- and 5-dimensional solutions is exclusively related to non-trivial aspects of Chern-Simons terms.

The off-shell dimensional reduction method can be used in many more situations. For instance, in performing the c-map of the topological string.

