

*Astron. Astrophys. Suppl. Ser.* 55, 171-177 (1984)

## Clean lines in the solar flux spectrum

R. J. Rutten and E. B. J. van der Zalm

Sterrewacht « Sonnenborgh », Zonnenburg 2, 3512 NL Utrecht, The Netherlands

*Received July 19, accepted October 12, 1983*

**Summary.** — We list profile parameters of 602 unblended lines in the Sacramento Peak Atlas of the visual solar irradiance spectrum, and we compare these to our earlier measurements of the same lines in the Jungfrauoch Atlas of the solar disk-center intensity spectrum.

**Key words :** solar spectroscopy — stellar abundances.

### 1. Introduction.

In our previous paper (Rutten and van der Zalm, 1984; henceforth paper I) we have selected 750 unblended lines present in the visual disk-center spectrum of the sun, and provide a list of various line-profile parameters. In this paper we extend these solar clean-line data by adding profile parameters measured from the full-disk spectrum. We so make an intermediate step between solar resolved-disk intensity spectrometry and stellar full-disk flux spectrometry; the list compiled here should be useful for stellar abundance analysis.

Paper I is based on the Jungfrauoch Atlas of the visual disk-center spectrum (Delbouille *et al.*, 1973; henceforth JJ Atlas). Here we use the Sacramento Peak Irradiance Atlas of the full-disk spectrum (Beckers *et al.*, 1976; henceforth SP Atlas). Each atlas has been recorded photoelectrically with a double-pass scanning-grating spectrometer, and each furnishes the highest spectral purity (spectral resolution, signal-to-noise ratio, absence of distortion) available for the respective spectrum. For our purposes their instrumental broadening is negligible; while the resolution of the SP Atlas is somewhat coarser, so is the inherent spectral detail of the flux spectrum. However, the appreciably larger noise level of the SP Atlas influences our results below.

### 2. Data reduction.

**2.1 BACKGROUND NORMALIZATION.** — We use the magnetic-tape editions of the two atlases and discuss only their region of overlap ( $\lambda\lambda 400.6-700.0$  nm). Both atlases suffer from scale variations which result from the record

concatenation necessary for segmented recording. This is illustrated in figure 1, where the comparison given in figure 5 of paper I is extended with the corresponding interval of the SP Atlas. The panels show the top 4% of each atlas, and also disk-center data taken with the broad-band Fourier Transform Spectrometer at Kitt Peak (Brault, 1978, example 2), which here supply a standard of quality. Each tracing is again normalized to a straight line (dotted) connecting the peaks at  $\lambda = 401.34$  nm and  $\lambda = 405.31$  nm, which are the highest in the FTS data and in the JJ Atlas, respectively. The upward spikes display the background windows in this crowded region on an exaggerated scale. The differences between their envelopes in the two lower panels are due to segment-by-segment tilts of the JJ Atlas records, of which the junctions are indicated. The SP Atlas (top) shows better background normalization, at least in the left half of the figure; but its peak-to-peak variations are larger, primarily due to the larger noise.

The varying segment tilts encumber locating the *true* continua  $I_c^T$  and  $F_c^T$ , whereas the noise decreases the precision with which one can locate the *local* continuum  $F_c^L$  (see paper I for notation and definitions; we use  $F$  for irradiance and call it flux). In section 2.1 of paper I we have worked out in detail how to convert measurements normalized to the local continuum into line parameters referring to the true continuum, i.e. corrected for the presence of unresolved blends. We refrain here from applying similar corrections to measurements from the SP Atlas because the errors in the evaluation of  $F_c^L$ , which are rather large due to the noise in the background flux values, exceed the corrections, which are small. We thus provide « local » line strengths and line depths only.

**2.2 LINE PARAMETERS.** — Of the 750 lines without resolved blends selected from the JJ Atlas in paper I, there are 638 present below the upper wavelength limit of the SP

Send offprint requests to : R. J. Rutten.

Atlas (700 nm). We have rejected 34 lines for which the adjacent maxima in the SP Atlas are so low that their halfwidths are ill-defined, and also  $\lambda 588.383$  and  $\lambda 697.047$ , which should have been rejected in paper I because they are blended by telluric H<sub>2</sub>O lines; the latter are stronger in the SP Atlas than in the JJ Atlas. For the remaining 602 lines we have measured the profile parameters listed in table I<sup>(1)</sup>.

The first column of table I specifies the wavelength of the line in nm, measured again as the normalized first moment of the lower third of the profile, on the wavelength scale of the SP Atlas. The identification (spectrum and multiplet number) is again copied from Moore *et al.* (1966). The parameter « Mode » is W for the lines for which the whole profile was used, B if only the blue half and R if only the red half was used, as in paper I. The halving serves to reduce the effects of unresolved blends often present in one line wing; we have halved the SP Atlas profiles following the choices made in paper I for the JJ Atlas profiles.

The measurement of the other profile parameters differs from paper I in the determination of the extent of a line and of the local continuum, because the noise in the SP Atlas bars the simple definition of a line as the interval between two successive maxima. Instead, we here define a line to begin and end at the first maxima of a running mean that are outside the extent of the line as it was measured in the JJ Atlas. The running mean is given by  $(F_{k-1} + 2F_k + F_{k+1})/4$ , where  $k$  is the sample counter of the successive tape-edition irradiance values  $F_k$ . We define the local continuum  $F_c^L$  equal to the highest of these two smoothed maxima, and measure the fractional line depth  $D_{SP}^L = (F_c^L - F_\lambda)/F_c^L$ , the full width at half maximum in pm  $\text{FWHM}_{SP}$ , and the logarithmic line strength  $\log(W_{SP}^L/\lambda)$  from this background level. The last column specifies the background, relative to the nominal continuum level of the SP Atlas  $F_c^{SP}$  (900 in tabular- and tape-edition units).

### 3. Discussion.

In figure 2 we show the differences between the strengths, depths and widths of the 602 lines in table I and the corresponding values given for the JJ Atlas in table IV of paper I, as functions of line strength (left) and line depth (right). Each panel shows a well-defined trend at right and large spread at left; we discuss these features separately.

What trends should we expect? Let us first consider the effects of the solar *rotation*. Simple theory (e.g. Unsöld, 1955, pp. 508-511; Gray, 1976, pp. 393-399) shows that if the intensity profile of a line, relative to its background, does not vary over the disk, then the flux profile is simply the convolution of the intensity profile with a rotation profile of which the shape is set by the continuum limb darkening. This convolution decreases the line depths and increases the line widths, but does not affect the line strengths. The changes increase with wavelength because the solar limb is less darkened at longer wavelengths, so that

the rotation profile is flatter: in the red, the west and east limbs contribute more Doppler-shifted signal than in the blue. We estimate the effect of rotational broadening from a simple numerical experiment. A plot of  $\text{FWHM}_{JJ}/\lambda$  against  $D_{JJ}^L$  (not shown) shows that disk-center lines of depth  $D^L = 0.5$  typically have  $\text{FWHM}/\lambda = 0.015 \pm 0.002$ . Smearing gaussians of this halfwidth with the solar rotation profile (with  $v \sin i = 1.9$  km/s) results in lowering of their amplitudes by  $10 \mp 3\%$  at  $\lambda = 550$  nm, with less than 1% amplitude change over the range  $\lambda \lambda 400$ -700 nm.

In addition, there are effects of center-to-limb variations in the *intensity profiles*. Most of these lines are better described by pure absorption than by scattering; they become shallower towards the limb, which results in reduction of their depths and strengths in the flux spectrum compared with the rotationally-broadened disk-center intensity spectrum. Furthermore, the broadening effects of granulation and waves, measured traditionally as micro- and macroturbulence, increase when viewed obliquely, resulting in decrease of the line depths, increase of the line widths, and, for lines on the flat part of the curve of growth, increase of the line strengths. These profile effects also increase with wavelength due to the decrease in limb darkening, but more steeply than the rotational broadening because the whole limb contributes, rather than its east and west parts only. The largest profile effect is generally the increase of the macroturbulence from 1 km/s at disk center to 2 km/s at the limb.

Let us now compare these expectations to the actual trends in figure 2. The mean vertical offsets of the stronger lines, best seen in the  $D_{JJ}^L$  panels at right, indeed display effects of smearing: the line widths (bottom) are larger in the flux spectrum and the depths (middle) are smaller, while the strengths (top) are about the same. The reductions in line depth are somewhat larger than the prediction for rotational broadening alone, especially at longer wavelengths. They show a pronounced increase with wavelength, which we attribute to the profile effects. The long-wavelength lines in the middle panels (crosses) show upturns for the strongest lines, i.e. smaller reduction in line depth; this decrease in the effect of the smearing on their line depth results from their line-core saturation. (A similar upturn is shown by the plot of  $\text{FWHM}_{JJ}/\lambda$  against  $D_{JJ}^L$  mentioned above.) The strongest blue lines present (circles) have not yet reached saturation and show no upturn.

We now turn to the *spread*. Some of it may be due to variations in the profile effects discussed above, but most of it results from measurement errors, primarily in the assignment of the local continuum  $F_c^L$  which is most sensitive to the noise in the SP Atlas. This is illustrated in the middle-right panel of figure 2, in which the increase of the spread for weak lines is most abrupt. (The fan-like structure at left is due to the three-digit discretization.) The many weak lines that quite erroneously have  $D_{SP}^L > D_{JJ}^L$  testify that  $J_c^L$  has often been overestimated, notwithstanding our smoothing. The size of the spread multiplied by  $D^L$  shows that the errors in  $J_c^L$  range up to 1%, in agreement with the general accuracy of 1% estimated for the SP Atlas by its authors; we find that heavier smoothing in the evaluation of  $J_c^L$  does not produce significant improvement.

<sup>(1)</sup> Copies on magnetic tape or punched cards can be obtained from E.v.d.Z.

The spread in the top panels is larger (0.1 dex = 26 %) because the equivalent widths are about twice as sensitive to errors in  $J_c^L$  as the line depths are. The many overestimations of  $J_c^L$  have caused the slight mean offsets shown by the weaker lines in the top panels.

Finally, we note that the spread in the top-left panel is two to three times smaller than the spread in the comparable bottom-left panel of figure 6 in paper I. The lines most suitable for abundance analysis, with  $-5.5 \leq \log(W/\lambda) \leq -5.2$ , have errors well within 0.1 dex, which is the best precision attainable using line

strengths in curve-of-growth methods (see Sect. 5 of paper I).

#### Acknowledgements.

We thank C. Zwaan and B. J. Oranje for comments.

*Note added in proof* : A detailed analysis of solar rotational broadening has been given in the PhD thesis of David H. Bruning, now at Mt Wilson.

#### References

- BECKERS, J. M., BRIDGES, Ch. A., GILLIAM, L. B. : 1976, *A High Resolution Spectral Atlas of the Solar Irradiance From 380 to 700 Nanometers*, Sacramento Peak Observatory (SP Atlas).
- BRAULT, J. W. : 1978, in « Future Solar Optical Observations : Needs and Constraints », eds. G. Godoli, G. Noci, A. Righini, *Osser. e Memorie dell' Oss. Astrof. di Arcetri* **106**, 33.
- DELBOUILLE, L., ROLAND, G., NEVEN, L. : 1973, *Photometric Atlas of the Solar Spectrum from  $\lambda 3000$  to  $\lambda 10000$* , Institut d'Astrophysique, Liège (JJ Atlas).
- GRAY, D. F. : 1976, *The Observation and Analysis of Stellar Photospheres*, Wiley Sons, New York.
- MOORE, Ch. E., MINNAERT, M. G. J., HOUTGAST, J. : 1966, « The Solar Spectrum 2935 Å to 8770 Å », *Nat. Bur. Stand. (U.S.) Monogr.* No. 61, Washington.
- RUTTEN, R. J., VAN DER ZALM, E. B. J. : 1984, *Astron. Astrophys. Suppl. Ser.* **55**, 143 (Paper I).
- UNSÖLD, A. : 1955, *Physik der Sternatmosphären*, Springer Verlag, Berlin.

TABLE I.

Wavelength	Ident	Mode	D <sup>L</sup>	FWHM	log(W/λ)	L <sub>c</sub> /L <sub>e</sub> <sup>L</sup>	L <sub>e</sub> <sup>L</sup> /L <sub>e</sub> <sup>SP</sup>	Wavelength	Ident	Mode	D <sup>L</sup>	FWHM	log(W/λ)	L <sub>c</sub> /L <sub>e</sub> <sup>L</sup>	L <sub>e</sub> <sup>L</sup> /L <sub>e</sub> <sup>SP</sup>
401.3233	Ti Ip	186	W	0.094	6.81	-5.739	0.993	480.8147	Fe I	633	W	0.306	8.04	-5.245	0.978
402.1335	Nd II	36	B	0.182	5.69	-5.799	0.986	480.9939	Fe I	579	B	0.213	8.41	-5.389	0.987
404.1907	Fe I	602	B	0.143	4.82	-5.732	0.952	481.3479	Co I	158	B	0.417	10.15	-5.018	0.984
404.2581	Ce II	140	B	0.152	4.81	-5.715	0.963	482.0413	Ti I	126	R	0.427	8.69	-5.080	0.985
405.0325	Zr II	43	B	0.305	6.74	-5.255	0.990	483.5874	Fe I	1068	B	0.497	8.96	-5.005	0.989
406.0268	Ti I	80	W	0.489	6.65	-5.045	0.976	483.6853	Cr I	144	B	0.159	8.44	-5.524	0.986
408.5723	Zr II	54	R	0.048	5.05	-6.208	0.956	483.7403	Ti Ip	250	B	0.028	8.20	-6.342	1.000
409.1559	Fe I	357	W	0.672	7.93	-4.833	0.971	488.0059	Cr I	167	B	0.069	9.22	-5.840	0.972
419.4493	Fe I	274	W	0.246	6.48	-5.394	0.979	488.5436	Fe I	966	R	0.660	9.75	-4.837	0.981
426.5926	Mn I	23	B	0.633	8.59	-4.861	0.975	489.2863	Fe I	1070	B	0.514	9.32	-4.960	0.976
426.6218	Ti I	252	R	0.098	6.38	-5.831	0.993	490.5138	Fe I	986	B	0.323	8.47	-5.195	0.976
428.1373	Ti I	44	B	0.333	7.27	-5.217	0.988	490.8608	Fe I	115	R	0.080	8.50	-5.786	0.983
428.1593	Fe Ip	171	B	0.105	6.01	-5.811	0.991	491.3976	Mn I	132	R	0.534	9.39	-4.926	0.981
431.9445	Fe I	214	B	0.205	6.71	-5.458	0.974	491.5236	Ti I	157	W	0.086	7.17	-5.849	0.986
434.9785	Ce II	59	W	0.085	6.55	-5.852	0.970	491.7234	Fe I	1066	B	0.577	9.80	-4.903	0.979
436.4858	Cr I	153	B	0.019	7.50	-6.649	0.983	492.6153	Ti I	39	R	0.079	8.69	-5.814	0.999
439.2308	Fe Ip	757	R	0.092	8.09	-5.687	0.981	493.5839	Mn I	177	R	0.565	9.72	-4.899	0.988
441.0523	Mn I	88	R	0.553	9.14	-4.873	0.975	493.6341	Cr I	166	B	0.446	8.98	-5.062	0.980
441.3403	Fe Ip	1046	B	0.086	12.49	-5.613	0.993	495.3214	Mn I	111	R	0.534	9.39	-4.930	0.980
443.9636	Fe I	515	W	0.247	7.42	-5.347	0.994	495.3719	Cr I	166	R	0.058	7.99	-5.975	0.997
443.9884	Fe I	116	W	0.572	8.07	-4.950	0.992	495.4293	Fe Ip	1093	B	0.017	3.33	-6.850	0.986
444.1273	Ti I	160	W	0.082	6.85	-5.890	0.974	496.1919	Fe I	845	B	0.270	8.70	-5.294	0.986
444.5475	Fe I	2	R	0.508	6.82	-5.069	0.975	496.2577	Fe I	1097	R	0.520	9.62	-4.935	0.986
444.6385	Nd II	49	B	0.149	5.61	-5.698	0.982	496.4719	Ti I	173	B	0.094	8.02	-5.793	0.988
446.5808	Ti I	146	W	0.446	8.09	-5.039	0.996	496.4931	Cr I	9	B	0.423	8.22	-5.130	0.974
447.0485	Mn I	86	B	0.717	10.43	-4.721	0.982	496.6813	Cr I	259	R	0.027	8.28	-6.340	0.986
448.5970	Fe Ip	825	W	0.237	7.79	-5.339	0.993	496.7528	Mn I	141	R	0.148	8.84	-5.550	0.972
449.1655	Cr I	95	R	0.256	7.86	-5.314	0.976	497.6698	Mn Ip	254	W	0.073	7.84	-5.895	0.984
449.1845	Cr I	83	W	0.127	6.81	-5.694	0.979	498.7623	Fe Ip	1094	W	0.037	7.73	-6.222	0.983
449.2308	Cr I	197	B	0.309	7.97	-5.210	0.994	499.2789	Fe I	1110	B	0.101	8.97	-5.708	0.983
450.4224	Fe Ip	988	B	0.030	13.76	-6.134	1.000	499.3353	Fe II	36	B	0.378	9.23	-5.110	0.988
451.2268	Ca I	24	R	0.263	9.19	-5.219	0.996	499.5411	Fe I	1113	R	0.151	8.53	-5.570	0.976
451.2740	Ti I	42	B	0.689	9.23	-4.786	0.993	499.5655	Mn I	145	B	0.200	7.95	-5.482	0.975
451.2995	Mn I	163	R	0.233	8.19	-5.319	0.997	500.3748	Mn I	50	B	0.370	8.35	-5.176	0.978
451.8027	Ti I	42	R	0.688	9.22	-4.804	0.999	500.4896	Mn I	20	B	0.151	8.99	-5.509	0.987
452.3080	Ce II	2	R	0.164	8.04	-5.468	0.980	501.0943	Mn I	144	R	0.487	9.13	-5.008	0.987
452.3404	Fe I	829	R	0.477	8.33	-5.014	0.986	501.6481	Fe I	1089	B	0.337	8.90	-5.180	0.989
454.3223	Fe Ip	893	R	0.049	8.11	-6.059	0.988	502.2874	Ti I	38	B	0.675	10.30	-4.820	0.979
454.6469	Fe I	1047	B	0.075	7.49	-5.898	0.983	502.6488	Mn I	158	B	0.031	7.08	-6.338	0.983
454.8770	Ti I	42	R	0.714	9.13	-4.800	0.988	504.7118	Fe Ip	1242	R	0.044	8.00	-6.165	0.973
455.1650	Fe I	972	B	0.317	7.44	-5.266	0.970	504.8854	Mn I	195	R	0.556	10.04	-4.898	0.968
455.6928	Fe I	638	R	0.300	7.72	-5.233	0.985	505.8503	Fe I	884	B	0.135	8.01	-5.660	0.994
456.0713	V I	109	B	0.111	7.19	-5.716	0.990	506.2103	Ti I	199	B	0.156	8.84	-5.544	0.991
456.2631	Ti I	7	R	0.147	7.12	-5.594	1.000	506.4068	Ti I	294	B	0.065	9.20	-5.893	0.996
456.3423	Ti I	266	R	0.127	7.38	-5.656	0.981	507.1488	Ti I	110	B	0.267	10.57	-5.228	0.977
456.6038	Fe Ip	1169	R	0.010	2.70	-7.168	1.000	509.4414	Mn I	164	R	0.317	9.28	-5.207	0.993
456.7409	Mn I	102	W	0.043	8.55	-6.083	0.999	511.2488	Cr I	19	B	0.046	6.73	-6.210	0.992
457.4219	Fe I	554	R	0.436	8.53	-5.047	0.976	511.3444	Ti I	109	B	0.284	8.33	-5.310	0.982
457.5109	Cr I	196	B	0.116	7.52	-5.706	0.971	511.9893	Fe Ip	960	R	0.051	9.31	-6.014	0.989
459.2055	Cr II	44	B	0.466	9.87	-4.964	0.984	512.2113	Cr I	19	B	0.127	7.45	-5.714	0.977
459.3530	Fe I	971	W	0.315	8.69	-5.179	0.991	514.1751	Fe I	114	B	0.711	12.29	-4.701	0.999
460.7086	Fe Ip	724	W	0.049	6.95	-6.105	0.988	514.5468	Ti I	109	B	0.373	9.00	-5.140	0.977
462.2745	Cr I	81	B	0.223	7.81	-5.395	0.967	514.5739	Fe Ip	931	R	0.059	9.91	-5.780	0.992
462.5923	Cr I	244	R	0.143	7.40	-5.610	0.984	515.2188	Ti I	4	R	0.389	8.50	-5.162	0.986
462.9073	Zr II	139	B	0.035	6.06	-6.325	0.990	515.7985	Mn I	111	B	0.202	8.55	-5.448	0.989
463.0785	Fe I	969	W	0.038	6.72	-6.249	0.989	515.9958	Fe Ip	1095	R	0.049	7.71	-6.127	0.993
463.3256	Cr I	186	W	0.110	7.55	-5.713	0.987	518.7918	Fe I	1032	R	0.494	9.40	-4.972	0.980
465.8298	Fe I	591	B	0.188	8.24	-5.460	0.972	519.6063	Fe I	1091	R	0.615	10.53	-4.850	0.974
467.0170	Fe II	25	R	0.309	8.26	-5.226	0.968	519.7165	Mn I	204	B	0.256	10.30	-5.251	0.988
467.2837	Fe Ip	40	B	0.341	9.03	-5.146	0.966	520.0173	Cr I	201	B	0.219	10.77	-5.295	0.990
469.0793	Ti I	76	B	0.043	5.26	-6.295	0.972	520.6806	Fe Ip	1095	R	0.071	9.49	-5.812	0.988
470.0611	Cr I	62	B	0.187	7.18	-5.525	0.966	521.3808	Fe I	962	B	0.076	8.98	-5.862	0.990
470.6309	Fe Ip	890	R	0.092	6.09	-5.876	0.980	521.4133	Cr I	193	B	0.182	8.54	-5.494	0.988
470.8016	Cr I	186	B	0.557	9.48	-4.917	0.966	521.9705	Ti I	4	R	0.279	8.51	-5.285	0.990
470.9713	Mn I	21	R	0.629	9.95	-4.838	0.970	522.3188	Fe I	880	B	0.300	9.11	-5.243	0.985
471.6833	Fe I	634	R	0.076	8.63	-5.793	0.980	523.0983	Ti I	215	W	0.020	8.54	-6.471	0.983
472.0129	Fe IIp	54	B	0.048	7.88	-6.090	0.983	523.4631	Fe II	49	W	0.649	11.66	-4.790	0.981
472.2163	Zn I	2	B	0.631	10.53	-4.771	0.989	523.7325	Cr II	43	W	0.446	10.82	-4.999	0.982
472.6138	Fe I	384	R	0.195	7.57	-5.474	0.974	523.8256	Fe I	962	W	0.032	10.32	-6.141	0.991
473.7353	Cr I	145	R	0.524	9.23	-4.959	0.972	523.8963	Cr I	59	B	0.176	8.64	-5.485	0.990
474.3826	Sc I	14	B	0.080	10.81	-5.715	0.989	524.0873	V I	131	W	0.050	10.92	-5.931	0.988
474.5135	Fe I	67	B	0.150	6.86	-5.650	0.966	524.1463	Cr I	59	W	0.043	8.86	-6.081	0.990
474.8733	La II	65	B	0.051	6.78	-6.077	0.987	524.3783	Fe I	1089	R	0.543	10.17	-4.936	0.980
475.4368	Co I	156	R	0.070	6.92	-5.973	0.967	524.7289	Ti I	183	B	0.087	6.72	-5.932	0.975
475.7850	Ru I	12	B	0.025	7.27	-6.408	0.985	525.0218	Fe I	1	R	0.647	9.12	-4.910	0.971
475.8124	Ti I	233	B	0.471	8.45	-5.047	0.985	525.3029	Fe Ip	113	B	0.205	8.88	-5.414	0.994
475.8409	Mn I	193	W	0.039	8.84	-6.094	0.988	525.4643	Co I	187	B	0.045	6.78	-6.161	0.976
475.9276	Ti I	233	R	0.502	8.50	-5.011	0.984	525.5516	Nd II	43	B	0.066	6.48	-6.077	0.966
476.0068	Fe I	384	R	0.078	6.81	-5.942	0.978	525.9739	Pr II	35	B	0.033	8.86	-6.209	0.994
476.7857	Cr I	231	B	0.194	7.87	-5.462	0.988	525.9973	Ti I	298	W	0.064	9.11	-5.931	0.989
477.0668															



TABLE I (continued).

Wavelength	Ident	Mode	$\frac{L}{D}$	FWHM	$\log(\frac{L}{\lambda})$	$\frac{L}{F_c} / \frac{f_p}{F_c}$	Wavelength	Ident	Mode	$\frac{L}{D}$	FWHM	$\log(\frac{L}{\lambda})$	$\frac{L}{F_c} / \frac{f_p}{F_c}$		
531.0693	Cr II	43	R	0.120	9.89	-5.608	0.989	567.0853	V I	36	R	0.150	12.16	-5.460	0.985
531.1638	Hf II	37	R	0.027	6.96	-6.503	0.979	567.7698	Fe Ip	1057	B	0.060	11.15	-5.917	0.984
531.2653	Co I	197	B	0.062	8.81	-5.958	0.983	567.8383	Fe Ip	982	R	0.043	10.44	-6.073	0.986
531.2858	Cr I	225	R	0.183	9.93	-5.429	0.983	567.9929	Ti I	269	B	0.044	9.73	-6.126	0.982
531.8353	Sc II	22	B	0.116	9.58	-5.634	0.988	568.0246	Fe I	1026	R	0.102	10.12	-5.699	0.990
531.8768	Cr I	225	W	0.153	9.12	-5.544	0.985	568.6836	Sc I	12	R	0.075	9.86	-5.869	0.979
532.0040	Fe I	877	B	0.191	9.22	-5.460	0.979	569.0429	Si I	10	B	0.388	12.40	-5.018	0.985
532.1111	Fe I	1165	R	0.382	10.04	-5.065	0.991	570.1105	Si I	10	B	0.295	12.17	-5.166	0.987
532.5283	Co I	192	W	0.084	9.43	-5.795	0.976	570.1553	Fe I	209	R	0.639	11.90	-4.834	0.983
532.5557	Fe II	49	R	0.383	10.29	-5.084	0.981	570.2323	Cr I	203	W	0.203	12.12	-5.331	0.994
532.6819	Fe I	1147	B	0.126	9.93	-5.617	0.978	570.2663	Ti I	249	B	0.070	10.58	-5.849	0.993
534.2707	Co I	190	B	0.293	9.95	-5.199	0.992	571.3898	Ti I	249	W	0.038	8.70	-6.198	0.992
534.4763	Cr I	225	R	0.085	9.05	-5.785	0.993	571.7838	Fe I	1107	B	0.490	11.59	-4.936	0.978
534.7719	Ni I	145	B	0.049	9.65	-6.033	0.990	571.9813	Cr I	119	B	0.044	7.09	-6.182	0.988
534.8323	Cr I	18	R	0.737	12.48	-4.710	0.996	572.5658	V I	135	W	0.018	11.45	-6.470	0.989
535.2048	Co I	172	W	0.198	10.67	-5.381	0.991	572.7653	V I	35	B	0.063	11.50	-5.886	0.990
535.6093	Sc I	17	R	0.020	12.23	-6.409	1.000	572.9203	Cr I	257	W	0.028	12.83	-6.228	0.996
535.7186	Sc II	30	B	0.048	8.51	-6.087	0.988	573.1770	Fe I	1087	R	0.479	10.99	-4.971	0.993
535.8123	Fe Ip	628	R	0.102	8.16	-5.720	0.988	573.7069	V I	35	B	0.081	12.86	-5.745	0.985
535.9203	Co I	194	W	0.099	9.33	-5.713	0.990	573.8238	Fe I	1084	B	0.115	10.99	-5.634	0.989
537.6834	Fe I	1132	B	0.154	8.76	-5.589	0.986	573.9473	Ti I	228	B	0.078	9.20	-5.868	0.991
537.7043	La II	95	B	0.038	5.87	-6.336	0.986	574.1853	Fe I	1086	W	0.280	10.38	-5.257	0.984
537.9579	Fe I	928	R	0.543	10.38	-4.919	0.992	574.2968	Fe I	1084	B	0.102	9.41	-5.755	0.982
538.0323	C I	11	B	0.129	14.19	-5.438	0.985	574.8359	Ni I	45	R	0.266	9.90	-5.281	0.994
538.6336	Fe I	1064	B	0.311	9.70	-5.212	0.988	575.2037	Fe I	1180	R	0.448	11.04	-5.020	0.990
538.9487	Fe I	1145	R	0.631	11.41	-4.822	0.973	575.4408	Fe I	866	W	0.102	8.54	-5.817	0.961
539.2333	Ni I	250	B	0.119	9.32	-5.659	0.983	575.9259	Fe I	1184	B	0.070	8.48	-5.959	0.983
539.5219	Fe I	1143	B	0.198	9.20	-5.432	0.988	576.0346	Fe I	867	B	0.213	9.39	-5.422	0.988
539.8288	Fe I	1145	R	0.586	10.92	-4.877	0.988	576.0835	Ni I	231	W	0.295	10.54	-5.227	0.976
540.1269	Fe I	1146	B	0.244	9.45	-5.325	0.992	576.6338	Ti I	309	R	0.081	9.21	-5.831	0.992
540.2782	Y II	35	B	0.122	9.31	-5.656	0.986	577.5081	Fe I	1087	R	0.478	11.27	-4.976	0.999
540.6778	Fe I	1148	R	0.343	9.88	-5.132	0.996	577.8458	Fe I	209	W	0.208	9.86	-5.407	0.994
541.2790	Fe Ip	1162	R	0.182	9.26	-5.475	0.982	578.3065	Cr I	188	R	0.276	11.13	-5.231	0.986
541.7041	Fe I	1148	B	0.319	9.74	-5.160	0.993	578.4659	Fe I	686	B	0.237	9.95	-5.353	0.984
542.2157	Fe Ip	1145	R	0.100	8.64	-5.773	0.980	578.5973	Ti I	309	W	0.099	10.58	-5.717	0.976
542.5252	Fe II	49	R	0.367	10.36	-5.097	0.986	579.3073	Si I	9	B	0.306	12.77	-5.139	0.982
542.6256	Ti I	3	B	0.068	9.01	-5.910	0.984	579.3916	Fe I	1086	R	0.295	10.54	-5.221	0.982
543.6299	Fe I	1161	B	0.366	10.17	-5.116	0.989	580.5223	Ni I	234	W	0.332	11.70	-5.132	0.989
543.8043	Fe I	1237	B	0.026	6.59	-6.322	0.988	580.6729	Fe I	1180	R	0.419	11.39	-5.040	0.985
543.8309	Ti I	108	W	0.036	9.61	-6.150	0.981	580.9883	Fe Ip	1084	B	0.021	12.63	-6.370	0.990
544.3403	Fe Ip	1059	B	0.039	10.61	-6.091	0.995	581.1915	Fe I	1022	B	0.103	9.50	-5.745	0.990
544.8913	Ti I	259	R	0.016	7.11	-6.703	0.988	581.2833	Ti I	309	B	0.021	11.19	-6.406	0.990
546.4289	Fe I	1030	R	0.355	9.88	-5.141	0.990	581.4808	Fe I	1086	W	0.199	10.87	-5.385	0.993
546.8106	Ni I	192	B	0.126	8.78	-5.629	0.991	581.9933	V II	99	B	0.030	13.58	-6.157	0.991
547.0093	Fe I	1144	B	0.235	10.30	-5.312	0.988	582.3178	Fe II	164	W	0.018	9.87	-6.496	0.988
547.1199	Ti I	106	B	0.085	10.19	-5.745	0.996	582.4413	Fe Ip	58	B	0.026	10.63	-6.320	0.991
547.3383	Y II	27	B	0.056	8.85	-6.076	0.972	582.7878	Fe Ip	552	R	0.105	10.46	-5.673	0.994
547.4225	Ti I	108	R	0.107	9.58	-5.712	0.983	583.2483	Ti I	309	W	0.019	11.26	-6.404	0.993
547.4459	Ti I	259	R	0.036	6.88	-6.301	0.986	583.5103	Fe Ip	1084	B	0.123	9.64	-5.661	0.983
548.4633	Sc I	16	B	0.030	11.61	-6.218	0.993	583.7708	Fe I	1129	B	0.087	10.54	-5.769	0.989
549.0153	Ti I	107	B	0.217	9.01	-5.423	0.981	583.8673	Cr I	119	R	0.038	9.93	-6.174	0.988
549.1836	Fe I	1031	R	0.123	9.81	-5.634	0.990	584.4593	Cr I	119	B	0.040	7.39	-6.264	0.986
549.4468	Fe I	1024	B	0.248	9.86	-5.308	0.978	584.4916	Fe I	1056	B	0.031	8.09	-6.318	0.990
550.3893	Ti I	287	B	0.133	7.88	-5.692	0.976	584.7000	Ni I	44	B	0.204	10.05	-5.421	0.993
552.0498	Sc I	15	B	0.072	9.50	-5.881	0.990	584.9693	Fe Ip	922	W	0.075	9.76	-5.869	0.992
552.1298	Fe Ip	1162	B	0.049	8.27	-6.110	0.972	585.2223	Fe I	1178	W	0.343	10.89	-5.131	0.995
552.2449	Fe I	1108	W	0.394	9.88	-5.112	0.984	585.3153	Fe I	95	B	0.069	9.66	-5.914	0.991
552.4245	Fe I	1059	R	0.042	9.85	-6.110	0.988	585.5083	Fe I	1179	B	0.194	10.94	-5.411	0.983
552.6815	Sc II	31	B	0.616	11.80	-4.837	0.986	585.6093	Fe I	1128	R	0.295	10.52	-5.226	0.992
553.7113	Ni I	188	B	0.031	10.25	-6.256	0.994	585.8788	Fe Ip	1084	R	0.118	9.96	-5.649	0.991
553.9281	Fe I	871	W	0.167	9.43	-5.495	0.983	585.9597	Fe I	1181	R	0.544	12.11	-4.898	0.984
554.6993	Fe I	1061	B	0.223	9.61	-5.367	0.977	586.1113	Fe Ip	1084	R	0.076	10.11	-5.832	0.989
555.2238	Sc II	25	R	0.044	7.34	-6.175	0.982	586.2363	Fe I	1180	W	0.592	13.08	-4.815	0.985
555.2698	Fe Ip	1281	B	0.072	10.39	-5.857	0.983	586.6457	Ti I	72	R	0.410	10.79	-5.076	0.993
556.0218	Fe I	1164	B	0.456	10.45	-5.020	0.992	586.7568	Ca I	46	B	0.211	11.97	-5.317	0.993
556.8870	Fe I	869	B	0.096	9.70	-5.789	0.982	587.6288	Fe Ip	1084	W	0.032	8.88	-6.316	0.975
557.7026	Fe I	1314	R	0.110	10.36	-5.652	0.993	587.9498	Fe Ip	1201	B	0.091	10.86	-5.709	0.990
557.8725	Ni I	47	R	0.484	10.34	-5.000	0.985	588.0033	Fe I	1201	W	0.105	10.44	-5.702	0.988
558.7578	Fe I	1026	B	0.325	10.03	-5.204	0.970	588.0289	Ti I	71	B	0.052	11.66	-6.002	0.986
558.7863	Ni I	70	R	0.499	10.22	-5.000	0.976	588.1285	Fe Ip	1178	R	0.129	10.05	-5.596	0.986
558.9363	Ni I	205	R	0.260	9.70	-5.289	0.994	588.4448	Cr I	119	B	0.023	8.91	-6.450	0.983
559.3741	Ni I	206	B	0.371	11.05	-5.070	0.989	590.2466	Fe I	1234	B	0.129	8.02	-5.723	0.993
559.5063	Fe Ip	1314	W	0.059	9.55	-5.970	0.983	590.3313	Ti I	71	B	0.050	10.58	-6.034	0.999
560.6993	Ni I	205	B	0.031	6.27	-6.416	0.990	590.5675	Fe I	1181	R	0.455	11.67	-4.982	0.994
560.8980	Fe Ip	1108	W	0.099	10.20	-5.665	0.995	590.6508	Ti I	105	B	0.043	10.79	-6.079	0.988
560.9966	Fe Ip	866	R	0.049	11.54	-6.004	0.992	592.2118	Ti I	72	W	0.181	10.41	-5.467	0.998
561.1356	Fe Ip	869	B	0.091	8.55	-5.823	0.982	592.7794	Fe I	1175	B	0.353	11.34	-5.124	0.992
561.8639	Fe I	1107	W	0.440	10.54	-5.044	0.984	592.9683	Fe I	1176	B	0.329	11.24	-5.175	0.985
561.9603	Fe I	1161	W	0.298	10.38	-5.211	0.986	593.0189	Fe I	1180	W	0.588	13.55	-4.814	0.985
562.2213	Si I	11	R	0.044	12.69	-5.955	0.99								

TABLE I (continued).

Wavelength	Ident	Mode	D <sup>L</sup>	FWHM	log(W/λ)	L <sub>ν</sub> <sup>L</sup> /F <sub>ν</sub> <sup>L</sup>	L <sub>ν</sub> <sup>R</sup> /F <sub>ν</sub> <sup>R</sup>	Wavelength	Ident	Mode	D <sup>L</sup>	FWHM	log(W/λ)	L <sub>ν</sub> <sup>L</sup> /F <sub>ν</sub> <sup>L</sup>	L <sub>ν</sub> <sup>R</sup> /F <sub>ν</sub> <sup>R</sup>
603.4033	Fe Ip	1142	W	0.072	10.16	-5.895	0.992	639.2543	Fe I	109	B	0.156	11.16	-5.545	0.994
603.9729	V I	34	W	0.115	9.89	-5.689	0.988	639.3613	Fe I	168	R	0.692	16.75	-4.690	0.983
605.3683	Mn I	247	R	0.159	11.31	-5.484	0.989	641.6923	Fe II	74	B	0.300	11.83	-5.207	0.992
605.4079	Fe I	1142	B	0.079	11.72	-5.797	0.990	642.9913	Co I	81	B	0.026	15.28	-6.209	0.997
605.6011	Fe I	1259	B	0.515	13.47	-4.890	0.990	643.2685	Fe II	40	B	0.302	12.28	-5.211	0.992
606.4633	Ti I	69	B	0.072	9.93	-5.878	0.988	643.6413	Fe I	1016	B	0.082	11.34	-5.806	0.991
606.5495	Fe I	207	B	0.696	15.98	-4.700	0.978	644.0933	Mn I	39	W	0.048	12.54	-5.987	0.997
606.7638	Si I	15	R	0.024	9.23	-6.419	0.992	645.2330	V I	48	R	0.057	10.62	-5.942	1.000
607.8496	Fe I	1259	R	0.518	13.25	-4.894	0.984	645.5008	Co I	174	B	0.085	15.66	-5.670	0.989
607.9015	Fe I	1176	R	0.363	11.50	-5.113	0.988	645.6390	Fe II	74	B	0.435	13.42	-5.003	0.992
608.1449	V I	34	B	0.117	10.71	-5.660	0.986	646.4673	Ca I	19	B	0.100	10.38	-5.779	0.991
608.4103	Fe II	46	B	0.168	9.95	-5.524	0.990	647.1673	Ca I	18	R	0.604	13.81	-4.843	0.989
608.6288	Mn I	249	R	0.332	11.73	-5.141	0.991	647.7853	Co I	174	R	0.030	11.16	-6.276	0.995
608.6683	Co I	165	B	0.022	13.45	-6.294	0.990	648.1875	Fe I	109	W	0.476	12.18	-5.012	0.987
608.9570	Fe I	1327	R	0.296	11.61	-5.193	0.992	648.2806	Ni I	66	W	0.317	12.32	-5.173	0.994
609.0216	V I	34	R	0.280	10.83	-5.262	0.991	649.8942	Fe I	13	W	0.378	10.73	-5.170	0.988
609.1173	Ti I	238	B	0.129	9.77	-5.654	0.991	649.9653	Ca I	18	R	0.577	13.56	-4.870	0.994
609.3153	Co I	37	B	0.067	13.66	-5.808	0.989	650.8853	Ca I	18	R	0.079	10.77	-5.802	0.994
609.3646	Fe I	1177	B	0.252	11.19	-5.307	0.990	651.6089	Fe II	40	R	0.386	12.84	-5.077	0.991
609.4376	Fe I	1177	R	0.161	11.11	-5.479	0.991	651.8376	Fe I	342	B	0.435	11.86	-5.065	0.978
609.6668	Fe I	959	R	0.302	11.41	-5.214	0.988	653.1423	V I	48	B	0.049	11.42	-6.058	1.000
609.8249	Fe Ip	1200	B	0.136	10.52	-5.609	0.988	653.2879	Ni I	64	B	0.131	11.33	-5.623	0.998
609.8669	Ti I	304	B	0.045	11.88	-6.063	0.986	653.3936	Fe I	1197	B	0.292	12.57	-5.217	0.996
611.1075	Mn I	230	R	0.275	11.88	-5.221	0.996	658.6319	Ni I	64	B	0.312	12.84	-5.180	0.990
611.1663	V I	34	W	0.083	13.36	-5.701	0.994	659.1329	Fe I	1229	B	0.079	13.79	-5.765	0.994
611.8109	Ni I	230	B	0.024	14.37	-6.332	0.991	659.3880	Fe I	168	R	0.551	13.93	-4.891	0.986
612.0253	Fe I	14	B	0.049	9.37	-6.077	0.988	659.8602	Ni I	249	R	0.188	13.54	-5.346	0.999
612.1013	Ti I	153	R	0.031	8.78	-6.340	0.984	659.9116	Ti I	49	B	0.077	10.97	-5.889	0.992
612.5023	Si I	30	W	0.207	13.59	-5.282	0.988	660.4596	Sc II	19	R	0.268	12.62	-5.234	0.998
612.6223	Ti I	69	R	0.195	10.18	-5.441	0.990	660.6963	Ti IIp	91	B	0.054	11.88	-6.008	0.983
612.8981	Ni I	42	W	0.231	10.12	-5.379	0.989	660.8033	Fe I	109	W	0.147	11.39	-5.546	0.997
613.0138	Ni I	248	B	0.184	10.48	-5.465	0.990	662.5023	Fe I	13	R	0.127	12.13	-5.596	0.995
613.3963	Ni I	229	W	0.044	11.22	-6.075	0.986	662.7549	Fe I	1174	B	0.207	12.28	-5.378	0.991
613.5368	V I	34	B	0.080	11.25	-5.812	0.979	663.0013	Cr I	16	R	0.052	12.66	-5.932	0.995
614.5016	Si I	29	R	0.256	13.95	-5.179	0.989	663.2443	Co I	111	R	0.057	12.85	-5.911	0.987
614.9248	Fe II	74	R	0.282	11.94	-5.229	0.987	663.3753	Fe I	1197	W	0.428	13.23	-5.036	0.965
615.1623	Fe I	62	W	0.420	11.06	-5.083	0.989	663.5128	Ni I	264	R	0.177	12.41	-5.451	0.985
615.5703	Si I	29	R	0.047	11.41	-6.024	0.982	666.1083	Cr I	282	B	0.087	13.44	-5.730	0.993
615.6803	O I	10	R	0.034	11.97	-6.135	0.990	666.1338	Ni I	246	R	0.051	10.75	-6.063	0.994
615.7731	Fe I	1015	B	0.471	11.83	-5.008	0.980	666.7433	Fe I	168	B	0.045	13.25	-6.026	0.992
615.9383	Fe I	1175	B	0.101	12.42	-5.667	0.987	666.7716	Fe I	1228	B	0.066	9.19	-6.008	0.988
616.0748	Na I	5	B	0.383	13.38	-5.044	0.983	668.7493	Y I	1	W	0.032	11.81	-6.210	1.000
616.5363	Fe I	1018	R	0.365	11.52	-5.120	0.985	669.8678	Al I	5	W	0.141	13.81	-5.488	0.994
616.6440	Ca I	20	R	0.504	12.67	-4.935	0.985	669.9143	Fe I	1228	B	0.061	10.52	-6.003	0.993
617.3343	Fe I	62	R	0.527	12.20	-4.922	0.986	670.3575	Fe I	268	R	0.286	12.08	-5.256	0.994
617.6818	Ni I	228	W	0.452	13.03	-4.967	0.988	670.4493	Fe I	1052	B	0.050	13.27	-5.957	0.999
617.7245	Ni I	58	W	0.128	9.87	-5.647	0.983	671.0323	Fe I	34	W	0.123	11.49	-5.651	0.992
618.6713	Ni I	229	B	0.236	11.42	-5.328	0.989	671.3743	Fe I	1255	R	0.159	12.25	-5.498	0.996
618.7408	Fe Ip	342	B	0.028	10.53	-6.334	0.984	672.5356	Fe I	1052	B	0.131	10.76	-5.626	0.991
618.7993	Fe I	959	R	0.368	12.11	-5.092	0.991	672.6673	Fe I	1197	B	0.329	13.07	-5.157	0.991
619.9506	Fe I	208	R	0.042	9.88	-6.106	0.992	673.3158	Fe I	1195	R	0.195	12.19	-5.408	0.999
620.0321	Fe I	207	B	0.542	12.61	-4.921	0.986	673.6513	Fe Ip	1122	B	0.015	8.44	-6.749	0.998
620.4613	Ni I	226	B	0.176	11.83	-5.448	0.989	673.9523	Fe I	34	R	0.095	11.16	-5.756	0.996
621.3438	Fe I	62	W	0.595	12.79	-4.871	0.992	674.3128	Ti I	48	R	0.137	12.81	-5.561	0.991
621.3873	V I	20	B	0.043	10.30	-6.084	0.993	674.5968	Fe I	1005	B	0.051	13.33	-5.960	0.992
621.6358	V I	19	W	0.262	12.93	-5.215	0.993	675.0158	Fe I	111	W	0.514	13.47	-4.954	0.993
622.0479	Ti I	293	R	0.075	9.83	-5.895	0.984	675.3463	Fe Ip	1196	W	0.050	12.02	-6.009	0.996
622.0783	Fe I	958	B	0.151	10.31	-5.580	0.983	675.6543	Fe Ip	1120	W	0.026	15.82	-6.242	0.996
622.3989	Ni I	228	B	0.222	11.44	-5.360	0.992	676.7785	Ni I	57	R	0.542	13.16	-4.927	0.991
622.4513	V I	20	B	0.051	10.19	-6.060	0.992	677.2323	Ni I	127	R	0.353	12.57	-5.136	0.992
622.6739	Fe I	981	R	0.237	11.70	-5.301	1.000	678.6863	Fe I	1052	W	0.188	12.28	-5.426	0.995
623.0093	Ni I	227	B	0.157	10.53	-5.538	0.982	679.3263	Fe I	1005	B	0.100	11.45	-5.741	0.996
623.2649	Fe I	816	W	0.557	13.60	-4.856	0.990	679.6128	Fe I	1007	W	0.078	13.16	-5.792	0.989
623.3193	V I	20	R	0.038	12.91	-6.071	0.995	679.8456	Ca I	31	B	0.038	11.00	-6.230	0.985
623.9353	Sc I	2	B	0.053	9.04	-6.117	0.987	680.1852	Fe Ip	34	W	0.016	9.40	-6.693	0.997
623.9943	Fe II	74	R	0.090	10.89	-5.802	0.979	680.4278	Fe I	1225	R	0.111	11.23	-5.689	0.992
624.0651	Fe I	64	R	0.409	11.31	-5.087	0.993	680.6850	Fe I	268	B	0.276	10.81	-5.326	0.999
624.2839	V I	19	B	0.073	13.50	-5.803	0.988	681.0266	Fe I	1197	R	0.349	13.31	-5.125	0.991
624.3113	V I	19	B	0.218	11.69	-5.372	0.983	681.4958	Co I	54	R	0.119	12.11	-5.598	0.996
624.5618	Sc II	28	B	0.265	11.69	-5.276	0.987	682.0375	Fe I	1197	B	0.281	13.43	-5.213	0.987
625.1833	V I	19	W	0.116	12.72	-5.579	0.998	682.4843	Fe Ip	1280	B	0.022	10.91	-6.467	0.994
625.2573	Fe I	169	R	0.687	15.37	-4.711	0.992	683.3238	Fe I	1194	R	0.072	11.76	-5.862	0.991
625.8111	Ti I	104	W	0.412	11.40	-5.091	0.991	683.7008	Fe I	1225	R	0.135	11.94	-5.584	0.991
626.5143	Fe I	62	W	0.599	13.22	-4.847	0.997	683.9843	Fe I	205	R	0.234	11.08	-5.373	0.988
627.0233	Fe I	342	R	0.429	11.27	-5.057	0.999	684.2048	Ni I	126	R	0.196	12.07	-5.417	0.989
627.4663	V I	19	W	0.070	11.18	-5.809	0.996	684.3663	Fe I	1173	B	0.413	13.66	-5.041	0.991
628.5168	V I	19	W	0.087	10.90	-5.790	1.000	684.8573	Si I	37	B	0.103	16.77	-5.540	0.998
629.0553	Fe Ip	208	B	0.040	12.43	-6.096	0.995	685.0443	Ni I	157	B	0.057	17.63	-5.796	0.998
630.0688	Sc II	28	B	0.052	12.06	-5.974	0.992	685.1632	Fe I	34	R	0.035	13.14	-6.193	0.999
630.3463	Fe I	1140</													

TABLE I (continued).

Wavelength	Ident	Mode	D	$L$	FWHM	$\log(W/\lambda)$	$L_{SP}^L/F_{\odot}^L$	$L_{JJ}^L/F_{\odot}^L$
698.8529	Fe I	167	W	0.254	12.54	-5.292	0.998	
699.6653	Ti I	256	B	0.018	9.93	-6.587	0.995	

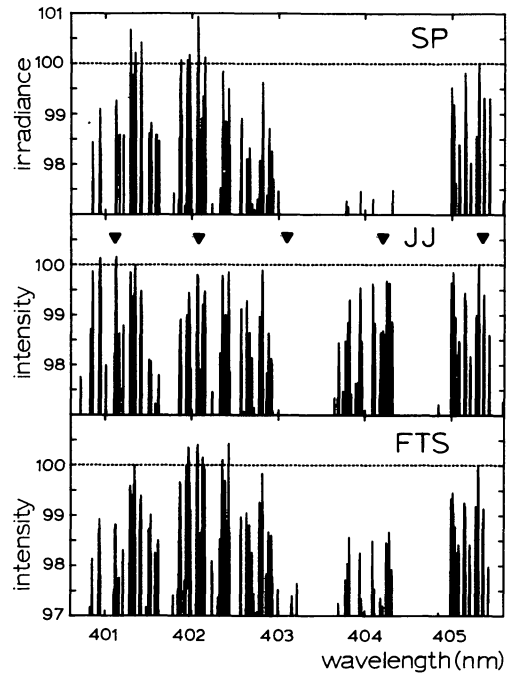


FIGURE 1. — Comparison of the background windows in the SP Atlas, in the JJ Atlas, and in KPNO-FTS data, for a short wavelength interval in the blue where these data sets overlap.

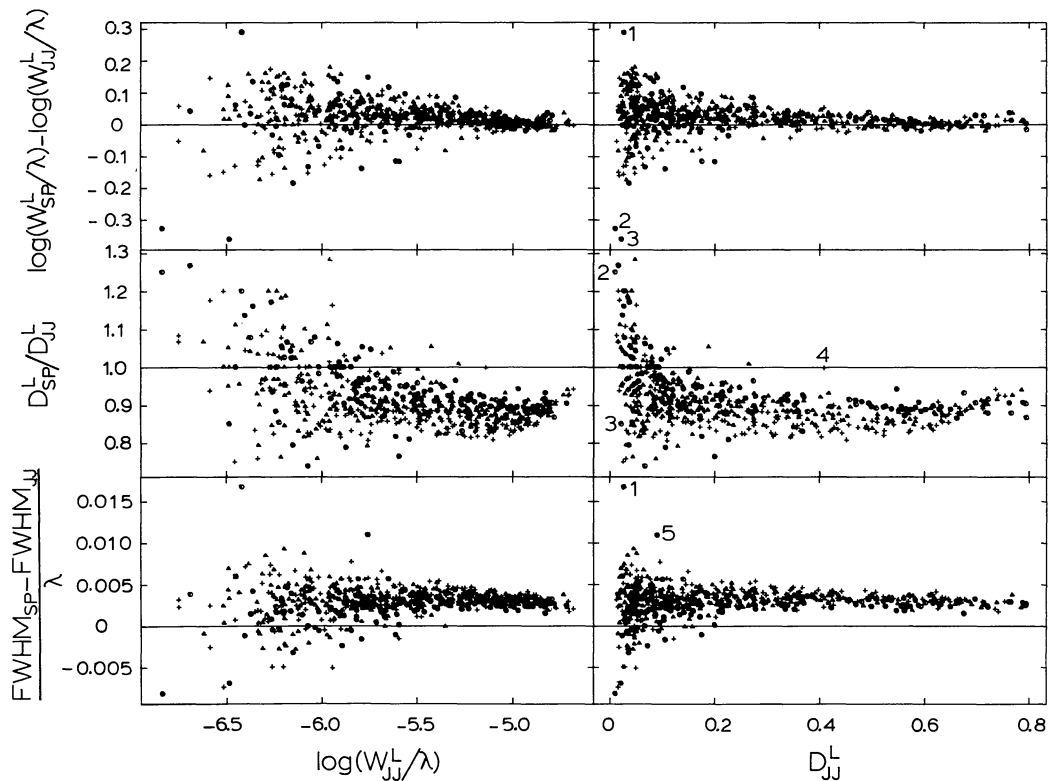


FIGURE 2. — Comparison of line-profile parameters measured from the SP Atlas and from the JJ Atlas for 602 clean lines. Horizontal : logarithmic line strengths from the JJ Atlas (left) ; line depths from the JJ Atlas (right). Vertical : difference in logarithmic line strengths (top) ; ratio of line depths (middle) ; difference in line halfwidths, normalized by the wavelength (bottom). All parameters are measured from the local continuum. Circles are lines from the wavelength interval  $\lambda\lambda 400.6-500$  nm, triangles from  $\lambda\lambda 500-600$  nm ; crosses from  $\lambda\lambda 600-700$  nm. (The numbers identify deviating lines, respectively : 1 =  $\lambda 450.421$ , 2 =  $\lambda 456.603$ , 3 =  $\lambda 495.430$ , 4 =  $\lambda 695.125$  and 5 =  $\lambda 441.339$ .)